# Lecture 4: Light + Color



#### Administrative

**HW0** is out, due 1/27

#### **Office Hours** and **Discussion Sections** start this week See <u>Google Calendar</u> for schedule and info

## Debugging

Learning to debug is an important goal for this class!

**Don't**: Debug via Piazza

Do:

- Google to find documentation for numpy / other libraries
- Figure out how numpy / other APIs work by poking around in an interactive shell, small standalone script, or notebook
- Try to isolate and reproduce errors in small standalone cases (unit testing!)

#### **Recap: Projection**



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#### Recap: Lenses

#### Pinhole Model



Mathematically correct Not quite correct in practice Reasonable approximation

#### **Reality: Lenses**



Necessary in practice Introduce complications Complications fixable



- A little bit about light and how you represent it
- A little bit about lighting and how it works

#### Your Very Own Camera



Where's the film/CCD?

January 26, 2021

Slide Credit: NIH

#### Your Very Own Camera



Where's the film/CCD?

Slide Credit: NIH



#### Demo Time

# https://bit.ly/2lNJ3xc



#### What is the Retina made of?



Slide Credit: J. Hays

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# Two Types of Photo Receptors

#### Cones

cone-shaped less sensitive operate in high light color vision

#### Rods

rod-shaped highly sensitive operate at night gray-scale vision



Slide Credit: J. Hays

# Rod / Cone Sensitivity



Slide Credit: J. Hays

#### Rod / Cone Distribution



Diagram Credit: B. A. Wandell, Foundations of Vision

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#### Electromagnetic Spectrum

![](_page_13_Figure_1.jpeg)

#### Why do we see light in these wavelengths?

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Slide Credit: J. Hays

#### The Physics of Light

![](_page_14_Figure_1.jpeg)

Slide Credit and Copyright: S. Palmer

## The Physics of Light

![](_page_15_Figure_1.jpeg)

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## The Physics of Light

![](_page_16_Figure_1.jpeg)

Slide Credit and Copyright: S. Palmer

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<u>Color wheel image</u> is free for commercial use under the <u>Pixabay license</u>

# The Physics of Light

![](_page_17_Figure_2.jpeg)

![](_page_17_Picture_3.jpeg)

Smooth transition from violet to red is an artifact of our visual system!

Slide Credit and Copyright: S. Palmer

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## Red-Green Color Blindness

"Peaks" of these red/green cones shifted, making it hard to distinguish red and green (%) 100 Μ RELATIVE ABSORBANCE S 50 400 450 500 550 600 650 WAVELENGTH (nm.)

Four possibilities:

- Deuteranomaly: Green cone shifted toward red
- Protanomaly: Red cone shifted toward green
- Deuteranopia: Green cone missing
- Protanopia: Red cone missing

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- Genes for red and green cones on X chromosome
- Men (XY): ~8% red-green color blind, (Northern European descent)
- Women (XX): ~0.5% red-green color blindness (Northern European descent)

# Red-Green Color Blindndess

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- Men (XY): ~8% red-green color blind, (Northern European descent)
- Women (XX): ~0.5% red-green color blindness (Northern European descent)
- Up to 15% of women may have four types of cones! (Tetrachromacy)

Jordan and Mollon, "A study of women heterozygous for colour deficiencies", Vision Research 1993

## Color Vision in Animals

#### Birds have four types of cones: can see ultraviolet light

![](_page_21_Figure_2.jpeg)

Some flowers have "Nectar Guides" visible under UV light

![](_page_21_Picture_4.jpeg)

Visible Light UV light

#### Human cones also sensitive to UV light! But blocked by lens

Image source: Wikipedia

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Image Source: Wikipedia

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#### Color Vision in Animals

Mantis Shrimp: Up to 16 types of photoreceptors! Can also detect polarization of light!

![](_page_22_Picture_2.jpeg)

Image source: Wikipedia

![](_page_22_Picture_5.jpeg)

#### **Artificial Cones**

![](_page_23_Picture_1.jpeg)

Slide Credit: S. Seitz

![](_page_23_Picture_4.jpeg)

#### Color Images

![](_page_24_Picture_1.jpeg)

Slide Credit: J. Hays

![](_page_24_Picture_4.jpeg)

## Color Images

#### Combined

![](_page_25_Picture_2.jpeg)

![](_page_25_Picture_3.jpeg)

![](_page_25_Picture_4.jpeg)

![](_page_25_Picture_5.jpeg)

![](_page_25_Picture_6.jpeg)

![](_page_25_Picture_7.jpeg)

Slide Credit: J. Hays

![](_page_25_Picture_10.jpeg)

#### Images in Python

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# Images in Python

Images are matrix / tensor im

**im[0,0,0]** top, left, red

im[y,x,c]
row y, column x, channel c

im[H-1,W-1,2]
bottom right blue

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Slide inspired by James Hays

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# 5 Things to Remember

- 1. Origin is top left
- 2. Rows are first index (what's the fastest direction for accessing?)
- Usually referred to as HWC (Height x Width x Channel) But you'll sometimes see CHW (especially with neural networks)
- 4. Typically stored as uint8 [0,255]
- 5. for y in range(H): for x in range(W): will run <u>1 million</u> <u>times</u> for a 1000x1000 image. A 4GHz processor can do only 4K clock cycles per pixel per second.

### Representing Colors

![](_page_29_Picture_1.jpeg)

#### Discussion time: how many numbers do you actually need for colored light? Assume all tuples (R,G,B) are legitimate colors (they are).

Image Credit: http://en.wikipedia.org/wiki/File:RGB\_illumination.jpg

### One Option: RGB

Pros

- 1. Simple
- 2. Common

<u>Cons</u> 1. Distances don't make sense 2. Correlated

![](_page_30_Picture_5.jpeg)

![](_page_30_Picture_6.jpeg)

G

B

![](_page_30_Picture_8.jpeg)

eren de la constante de la constan

Slide Credit: J. Hays, RGB cube: https://en.wikipedia.org/wiki/RGB\_color\_model

![](_page_30_Picture_12.jpeg)

# Another Option: HSV

<u>Pros</u> 1. Intuitive for picking colors 2. Sort of common

3. Fast to convert

![](_page_31_Picture_3.jpeg)

<u>Cons</u> 1. Not as good as other better spaces

![](_page_31_Picture_5.jpeg)

Slide Credit: J. Hays, HSV cylinder: https://en.wikipedia.org/wiki/HSL\_and\_HSV

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## Another Option: YUV / YCbCr

<u>Pros</u> 1. Great for transmission / compression

Y = 0

<u>Cons</u> 1. Not as good as other better smart color spaces

Y = 0.5

![](_page_32_Picture_3.jpeg)

● (Cb=0.5, Cr=0.5)

![](_page_32_Picture_5.jpeg)

**Cb** (Y=0.5, Cr=0.5)

![](_page_32_Picture_7.jpeg)

**Cr** (Y=0.5,

Cb=05)

Slide Credit: J. Hays, YUV cube: https://en.wikipedia.org/wiki/YUV

## Another Option: LAB

<u>Pros</u> 1. Distances correspond with human judgment 2. Safe <u>Cons</u> 1. Complex to calculate (don't write it yourself, lots of fp calculations)

![](_page_33_Picture_3.jpeg)

(a=0,b=0)

![](_page_33_Picture_5.jpeg)

**a** (L=65,b=0)

![](_page_33_Picture_7.jpeg)

**b** (L=65,a=0)

Slide Credit: J. Hays, Lab diagram cube: https://en.wikipedia.org/wiki/CIELAB\_color\_space

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# Why are there so many?

- Each serves different functions
  - RGB: sort of intuitive, standard, everywhere
  - HSV: good for picking, fast to compute
  - YCbCr/YUV: fast to compute, compresses well
  - Lab: the right(?) thing to do, but "slow" to compute
- Pick based on what you need and don't sweat it: color really isn't crucial

## Other Types of "Images"

- Almost all of this class is about ordinary RGB images because this has driven a lot of applications
- However, there are lots of other images

#### Depth Map

![](_page_36_Picture_1.jpeg)

![](_page_36_Picture_3.jpeg)

#### Surface Normals

![](_page_37_Picture_1.jpeg)

![](_page_37_Picture_2.jpeg)

![](_page_37_Picture_3.jpeg)

![](_page_37_Picture_4.jpeg)

Legend

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#### **Science Data**

Magnetic Field in: x, y, z

Light at 9 ~wavelenths: 9.4nm, 13.1nm, 17.1nm 19.3nm, 21.1nm, 30.4nm 33.5nm, 160nm, 170nm

NASA Solar Dynamics Observatory observing solar flare

![](_page_38_Picture_4.jpeg)

![](_page_38_Picture_5.jpeg)

#### Volumes

Volumes: images with more dimensions.

Emerge in 3D reconstruction, medical imaging, temporal data

![](_page_39_Picture_3.jpeg)

From: Girdhar et al., *Learning a predictable and generative vector representation for objects*. ECCV 2016

## Other Types of "Images"

- Almost all of this class is about ordinary RGB images because this has driven a lot of applications
- However, there are lots of other images

#### So Far:

#### How do we represent light and its storage on film?

![](_page_41_Picture_2.jpeg)

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![](_page_42_Picture_0.jpeg)

![](_page_42_Picture_2.jpeg)

![](_page_43_Figure_1.jpeg)

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![](_page_44_Figure_1.jpeg)

#### What happens when light hits a surface? 1. Absorbed It's absorbed and converted into some other form of energy (e.g., a black shirt getting hot in the sun)

![](_page_45_Picture_1.jpeg)

# What happens when light hits a surface?

#### 2. Transmitted

Possibly bouncing around before going through or out (e.g. lenses bend and go through, milk bounces around)

![](_page_46_Picture_1.jpeg)

# What happens when light hits a surface?

#### 3. Reflected

It's reflected back, in one or more directions with varying amounts (e.g., mirror, or a white surface)

![](_page_47_Picture_1.jpeg)

What happens when light hits a surface?

**4. Everything** All of the above! Real surfaces often have combinations of all of these options.

# Modeling Light and Surfaces

![](_page_48_Figure_1.jpeg)

**Opaque Reflections** 

#### Bi-directional reflectance distribution function (BRDF): % reflected given <u>i</u>ncident angle to light reflected angle to the viewer.

Note: have not specified form of function.

## Specular and Diffuse Reflectionv

# Same lighting, as close as possible camera settings, but different **location**

![](_page_49_Picture_2.jpeg)

![](_page_49_Picture_4.jpeg)

#### Specular and Diffuse Reflection

![](_page_50_Picture_1.jpeg)

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### **Diffuse Reflection**

![](_page_51_Figure_1.jpeg)

#### Lambertian Surface

Light depends **only** on orientation of surface  $\phi_i, \theta_i$  to light. Result of random small facets. Looks identical at all views.

### **Diffuse Reflection**

![](_page_52_Figure_1.jpeg)

Lambert's Law N: surface normal S: source direction **and** strength ρ: how much is reflected

$$B = \rho N \cdot S$$
$$B = \rho \|S\| \cos(\theta)$$

### Specular Reflection

![](_page_53_Picture_1.jpeg)

**Specular Surface** 

Light reflected like a mirror, but spreads out in a "lobe" around the reflection ray

![](_page_53_Picture_5.jpeg)

### Specular Reflection

![](_page_54_Figure_1.jpeg)

#### <u>Phong Model</u> V: vector to viewer R: reflection ray α: shininess constant

 $B = (V^T R)^{\alpha}$ 

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#### BRDFs can be very complex

![](_page_55_Picture_1.jpeg)

![](_page_55_Picture_2.jpeg)

![](_page_55_Picture_3.jpeg)

![](_page_55_Picture_4.jpeg)

![](_page_55_Picture_5.jpeg)

![](_page_55_Picture_6.jpeg)

![](_page_55_Picture_7.jpeg)

![](_page_55_Picture_8.jpeg)

Slide Credit: L. Lazebnik

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![](_page_56_Figure_0.jpeg)

![](_page_56_Figure_1.jpeg)

## Shape from Shading

![](_page_57_Figure_1.jpeg)

- System of equations that's underdetermined (N equations, 2N unknowns, N+3 known)
  - **Solution**: Add more equations that enforce smoothness or finding a single surface.

#### Realistic Shape from Shading

![](_page_58_Figure_1.jpeg)

- System of equations that's underdetermined (N equations, 2N+3 unknowns)
  - **Solution**: need prior beliefs to disambiguate.

#### Ambiguity

![](_page_59_Picture_1.jpeg)

![](_page_59_Picture_2.jpeg)

![](_page_59_Picture_4.jpeg)

### Ambiguity

# Humans assume light from above (and the blueness also tells you distance)

![](_page_60_Picture_2.jpeg)

Photo Credit: https://en.wikipedia.org/wiki/Meteor\_Crater

![](_page_60_Picture_5.jpeg)

### Shape from Shading in Practice

https://www.youtube.com/watch?v=4GiLAOtjHNo

![](_page_61_Picture_3.jpeg)

### Modeling Light and Surfaces

![](_page_62_Figure_1.jpeg)

#### Color that reaches eye depends on color of light and color of material

![](_page_62_Picture_4.jpeg)

## Color Ambiguity: "The Dress"

Blue / Black dress under yellow light?

![](_page_63_Picture_2.jpeg)

White / Gold dress under blue light?

https://en.wikipedia.org/wiki/The\_dress

![](_page_63_Picture_6.jpeg)

#### Recap

![](_page_64_Figure_1.jpeg)

**Light and Surfaces** 

![](_page_64_Figure_3.jpeg)

**Color Vision** 

![](_page_64_Figure_5.jpeg)

#### **Color Spaces**

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Next Lecture: Math Review

![](_page_65_Picture_2.jpeg)