

# Bayesian sparse image reconstruction: Application to MRFM

Nicolas Dobigeon, Alfred O. Hero and Jean-Yves Tournet

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**Index Terms**—Deconvolution, MRFM imagery, sparse representation, Bayesian inference, MCMC methods.

## I. INTRODUCTION

For several decades, image deconvolution has received increasing interest in the literature [1], [2]. Deconvolution mainly consists of reconstructing images from observations provided by optical devices and may include denoising, deblurring or restoration. The applications are numerous including astronomy [3], medical imagery [4], remote sensing [5] and photography [6]. More recently, a new imaging technology, so-called Magnetic Resonance Force Microscopy (MRFM), has been developed (see [7] and [8] for recent reviews). This non-destructive method allows one to improve the detection sensitivity of standard magnetic resonance imaging [9]. Because of their potential atomic-level resolution<sup>1</sup>, the 2-dimensional or 3-dimensional images resulting from this technology are

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characterized by their sparsity. Indeed, as the observed objects are molecules, most of the image is empty space. In this paper, a hierarchical Bayesian model is proposed to perform reconstruction of such images.

Deconvolution of sparse signals or images has motivated research in spectral analysis in astronomy [10], seismic signal analysis in geophysics [11], [12], or deconvolution of ultrasonic B-scans [13], among other examples. We propose here a Bayesian model that is based on an appropriate prior distribution for the unknown image. This prior is composed of a weighted mixture of a standard exponential distribution and a mass at zero. When the non-zero part of this prior is chosen to be a centered normal distribution, this prior reduces to a Bernoulli-Gaussian process. This distribution has been widely used in the literature to build Bayesian estimators for sparse deconvolution problems (see [14]–[18] or more recently [19] and [20]). However, choosing a distribution with heavier tail may improve the sparsity inducement of the prior. Combining a Laplacian distribution with an atom at zero results in the LAZE prior. This distribution has been used in [21] to solve a denoising problem in a non-Bayesian quasi-maximum likelihood estimation framework. In [22], [23], this prior has also been used for sparse reconstruction of noisy images. In this paper, a new prior composed of a mass at zero and a single-sided exponential distribution is introduced. The main motivation of choosing this prior is to take into account the positivity and the sparsity of the pixels in the image. The full Bayesian posterior can then be derived from samples generated by Markov chain Monte Carlo (MCMC) methods [24].

With the prior modeling introduced above, the results of the sparse reconstruction critically depend on the parameters chosen to define the mixture. Unfortunately, estimating the “hyperparameters” involved in the prior distribution described above is a difficult task. Empirical solutions have been proposed in [22], [23] to deal with this issue. When compared with other standard methods, the results in [22], [23] are satisfactory at low signal to noise ratios (SNR). At high SNRs, these methods display increasingly biased estimation of the hyperparameters that can lead to unstable results. In the Bayesian estimation framework, two approaches are available to estimate these hyperparameters. One approach couples MCMC methods to an expectation-maximization (EM) algorithm or to a stochastic EM algorithm [25], [26] to maximize a penalized likelihood function. The second approach defines non-informative prior distributions for the hyperparameters; introducing a second level of hierarchy in the Bayesian formulation. This fully Bayesian approach, adopted in this paper, has been suc-

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characterized by their sparsity. Indeed, as the observed objects are molecules, most of the image is empty space. In this paper, a hierarchical Bayesian model is proposed to perform reconstruction of such images.

Deconvolution of sparse signals or images has motivated research for spectral analysis in astronomy [10], for seismic signal analysis in geophysics [11], [12] or for deconvolution of ultrasonic B-scans [13], among other examples. We propose here a fully Bayesian model that is based on an appropriate prior distribution for the unknown image. This prior is composed of a weighted mixture of a standard exponential distribution and a mass at zero. When the non-zero part of this prior is chosen to be a centered normal distribution, this prior reduces to a Bernoulli-Gaussian process. This distribution has been widely used in the literature to build Bayesian estimators for sparse deconvolution problems (see [14]–[18] or more recently [19] and [20]). However, choosing a distribution with heavier tail may improve the sparsity inducement of the prior. Combining a Laplacian distribution with an atom at zero results in the  $\ell_{37}$ ZE prior. This distribution has been used in [21] to solve a deconvolution problem in a non-Bayesian quasi-maximum likelihood estimation framework. In [22], [23], this prior has also been used for sparse reconstruction of noisy images. In this paper, a new prior composed of a mass at zero and a single-sided exponential distribution is introduced. The main motivation of choosing this prior is to take into account the positivity and the sparsity of the pixels in the image. The full Bayesian posterior can then be derived from samples generated by Markov chain Monte Carlo (MCMC) methods [24].

With the prior modeling introduced above, the results of the sparse reconstruction critically depend on the parameters chosen to define the mixture. Unfortunately, estimating the hyperparameters involved in the prior distribution described above is a difficult task. Empirical solutions have been proposed in [22], [23] to deal with this issue. When compared with other standard methods, the results in [22], [23] are satisfactory at low signal to noise ratios (SNR). At high SNRs, these methods display increasingly biased estimation of the hyperparameters that can lead to unstable results. In the Bayesian estimation framework, two approaches are available to estimate these hyperparameters. One approach couples MCMC methods to an expectation-maximization (EM) algorithm or to a stochastic EM algorithm [25], [26] to maximize a penalized likelihood function. The second approach defines non-informative prior distributions for the hyperparameters; introducing a second level of hierarchy in the Bayesian formulation. This Bayesian approach, adopted in this paper, has been suc-

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e.g. by maximizing the estimated posterior distribution.
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collected from a tobacco virus sample using a prototype MRFM instrument.
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, including MRFM. The principal weakness of these previous approaches is the sensitivity to hyperparameters that determine the prior distribution, e.g. the LAZE mixture coefficient and the weighting of the prior vs the likelihood function. The hierarchical Bayesian approach proposed in this paper circumvents these difficulties.
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Conjugate priors on the hyperparameters of the image prior are introduced. It is this step that makes our approach hierarchical Bayesian.
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and poor estimation leads to instability
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Bayes (EB) and Stein unbiased risk (SURE)
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However, instability was observed especially at higher signal-to-noise ratios (SNR).
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cessfully applied to signal segmentation [27]–[29] and semi-supervised unmixing of hyperspectral imagery [30].

In this paper, the response of the MRFM imaging device is assumed to be known. This standard assumption makes the sparse image reconstruction a non-blind deconvolution problem that is a standard linear inverse problem [31]. The hierarchical Bayesian formulation proposed here naturally introduces an appropriate regularization for the ill-posed problem where the hyperparameters are estimated in an unsupervised scheme. Only a few works in the literature [47] are dedicated to reconstruction of MRFM image data [32]–[35]. In [36], several techniques based on linear filtering [48] maximum-likelihood principle [9] have been proposed. [10] nevertheless, none of these models and algorithms takes advantage of the sparse nature of the image to be analyzed. More recently, Ting *et al.* has introduced sparsity penalized reconstruction methods [12] motivated by MRFM applications [23]. The reconstruction problem is [15] composed into a deconvolution step and a denoising step, yielding an iterative thresholding framework. However, in [17]–[16], the hyperparameters are estimated via a heuristic manner by applying [18] Stein’s unbiased risk estimator [19]. [20] Contrary to our fully Bayesian approach that allows them to be marginalized. As it has been pointed out above, this *ad hoc* hyperparameter choice can lead to unreliable results. Moreover, a [23] posterior analysis is not possible with the strategy proposed in [23]. As a consequence, the parameter estimation is only based on the peak of the penalized likelihood function, whose research thanks to the EM algorithm can be subjected to slow convergence and local maxima [38]. [24]

This paper is organized as follows. The deconvolution problem is formulated in Section II. The hierarchical Bayesian model [26] described in Section III. Section IV presents a Gibbs sampler that allows one to generate samples distributed according to the posterior of interest. [27] The simulation results, including comparison of performances, are presented in Section V for MRFM. Our main conclusions are reported in Section VII.

## II. PROBLEM FORMULATION

Let  $\mathbf{X}$  denote a  $l_1 \times \dots \times l_n$  unknown  $n$ -dimensional pixelated image to be recovered (e.g.  $n = 2$  or  $n = 3$ ). This image is available as a collection of  $P$  projections  $\mathbf{y} = [y_1, \dots, y_P]^T$  which follows the model:

$$\mathbf{y} = T(\boldsymbol{\kappa}, \mathbf{X}) + \mathbf{n}, \quad (1)$$

where  $T(\cdot, \cdot)$  stands for a bilinear function,  $\mathbf{n}$  is a  $P \times 1$  dimension noise vector and  $\boldsymbol{\kappa}$  is the kernel that characterizes the response of the imaging device. Typical point spread responses  $\boldsymbol{\kappa}$  of MRFM tip can be found in [39] for horizontal and vertical configurations. In (1),  $\mathbf{n}$  is an additive Gaussian noise sequence distributed according to  $\mathbf{n} \sim \mathcal{N}(\mathbf{0}, \sigma^2 \mathbf{I}_P)$ .

Note that in standard deblurring problems, the function  $T(\cdot, \cdot)$  represents the standard  $n$ -dimensional convolution operator  $\otimes$ . In this case, the image  $\mathbf{X}$  can be vectorized yielding the unknown image  $\mathbf{x} \in \mathbb{R}^M$  with  $M = P = l_1 l_2 \dots l_n$ . With this notation, Eq. (1) can be rewritten:

$$\mathbf{y} = \mathbf{H}\mathbf{x} + \mathbf{n} \quad \text{or} \quad \mathbf{Y} = \boldsymbol{\kappa} \otimes \mathbf{X} + \mathbf{N} \quad (2)$$

where  $\mathbf{y}$  (resp.  $\mathbf{n}$ ) stands for the vectorized version of  $\mathbf{Y}$  (resp.  $\mathbf{N}$ ) and  $\mathbf{H}$  is an  $P \times M$  matrix that describes convolution by the psf  $\boldsymbol{\kappa}$ .

The problem addressed in the following sections consists of estimating  $\mathbf{x}$  under sparsity and positivity constraints on  $\mathbf{x}$  given the observations  $\mathbf{y}$ , the psf  $\boldsymbol{\kappa}$  and the bilinear function<sup>2</sup>  $T(\cdot, \cdot)$ .

## III. HIERARCHICAL BAYESIAN MODEL

### A. Likelihood function

The observation model defined in (1) and the Gaussian properties of the noise sequence  $\mathbf{n}$  yield:

$$f(\mathbf{y}|\mathbf{x}, \sigma^2) = \left( \frac{1}{2\pi\sigma^2} \right)^P \exp\left( -\frac{\|\mathbf{y} - T(\boldsymbol{\kappa}, \mathbf{x})\|^2}{2\sigma^2} \right), \quad (3)$$

where  $\|\cdot\|$  denotes the standard  $\ell_2$  norm:  $\|\mathbf{x}\|^2 = \mathbf{x}^T \mathbf{x}$ .

### B. Parameter prior distributions

The unknown parameter vector associated with the observation model defined in (1) is  $\boldsymbol{\theta} = \{\mathbf{x}, \sigma^2\}$ . In this section, we introduce prior distributions for these two parameters; which are assumed to be independent.

1) *Image prior*: First let consider the exponential distribution with shape parameter  $a > 0$ :

$$g_a(x_i) = \frac{1}{a} \exp\left(-\frac{x_i}{a}\right) \mathbf{1}_{\mathbb{R}_+^*}(x_i), \quad (4)$$

where  $\mathbf{1}_{\mathbb{E}}(x)$  is the indicator function defined on  $\mathbb{E}$ :

$$\mathbf{1}_{\mathbb{E}}(x) = \begin{cases} 1, & \text{if } x \in \mathbb{E}, \\ 0, & \text{otherwise.} \end{cases} \quad (5)$$

Choosing  $g_a(\cdot)$  as prior distributions for  $x_i$  ( $i = 1, \dots, M$ ) leads to a MAP estimator of  $\mathbf{x}$  that corresponds to a maximum  $\ell_1$ -penalized likelihood estimate with a positivity constraint<sup>3</sup>. Indeed, assuming the component  $x_i$  ( $i = 1, \dots, P$ ) a priori independent allows one to write the full prior distribution for  $\mathbf{x} = [x_1, \dots, x_M]^T$ :

$$g_a(\mathbf{x}) = \left( \frac{1}{a} \right)^M \exp\left( -\frac{\|\mathbf{x}\|_1}{a} \right) \mathbf{1}_{\{\mathbf{x} \succ 0\}}(\mathbf{x}), \quad (6)$$

where  $\{\mathbf{x} \succ 0\} = \{\mathbf{x} \in \mathbb{R}^M; x_i > 0, \forall i = 1, \dots, M\}$  and  $\|\cdot\|_1$  is the standard  $\ell_1$  norm  $\|\mathbf{x}\|_1 = \sum_i |x_i|$ . This estimator has shown interesting sparse properties for Bayesian estimation [41] and signal representation [42].

Coupling a standard probability density function (pdf) with an atom at zero is another classical alternative to ensure sparsity. This strategy has for instance been used for located event detection [14] such as spike train deconvolution [11], [17]. In order to increase the sparsity of the prior, we propose

<sup>2</sup>In the following, for sake of conciseness, the same notation  $T(\cdot, \cdot)$  will be adopted for the bilinear operations used on  $n$ -dimensional images  $\mathbf{X}$  and used on  $M \times 1$  vectorized images  $\mathbf{x}$ .

<sup>3</sup>Note that a similar estimator using a Laplacian prior for  $x_i$  ( $i = 1, \dots, M$ ) was proposed in [40] for regression problems and is usually referred to as the LASSO estimator but without positivity constraint.

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	While it may be possible to extend our methods to unknown point spread functions, e.g., along the lines of [54], the case of sparse blind deconvolution is outside of the scope of this paper.		
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	asymptotically generates Bayes-optimal estimates of all image parameters, including the hyperparameters.		
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	Despite promising results, especially at low SNR, penalized likelihood approaches require iterative algorithms that are often slow to converge and can get stuck on local maxima [38]. In contrast to [23], the fully Bayesian approach presented in this paper converges quickly and produces estimates of the entire posterior and not just local maxima.		
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cessfully applied to signal segmentation [27]–[29] and semi-supervised unmixing of hyperspectral imagery [30].

In this paper, the response of the MRFM imaging device is assumed to be known. This standard assumption makes the sparse image reconstruction a non-blind deconvolution problem that is a standard linear inverse problem [31]. The hierarchical Bayesian formulation proposed here naturally introduces an appropriate regularization for the ill-posed problem where the hyperparameters are estimated in an unsupervised scheme. Only a few works in the literature are dedicated to reconstruction of MRFM image data [32]–[35]. In [36], several techniques based on linear filtering or maximum-likelihood principle have been proposed. Nevertheless, none of these models and algorithms takes advantage of the sparse nature of the image to be analyzed. More recently, Ting *et al.* has introduced sparsity penalized reconstruction methods motivated by MRFM applications [23]. The reconstruction problem is decomposed into a deconvolution step and a denoising step, yielding an iterative thresholding framework. However, in [23], the hyperparameters are estimated via a heuristic manner by applying the Stein’s unbiased risk estimator [37], contrary to our fully Bayesian approach that allows them to be marginalized. As it has been pointed out above, this *ad hoc* hyperparameter choice can lead to unreliable results. Moreover, a full posterior analysis is not possible with the strategy proposed in [23]. As a consequence, the parameter estimation is only based on the peak of the penalized likelihood function, whose research thanks to the EM algorithm can be subjected to slow convergence and local maxima [38].

This paper is organized as follows. The deconvolution problem is formulated in Section II. The hierarchical Bayesian model are described in Section III. Section IV presents a Gibbs sampler that allows one to generate samples distributed according to the posterior of interest. Some simulation results, including [30] comparison of performances, are presented in Section V for MRFM. [31] main conclusions are reported in Section VII.

## II. PROBLEM FORMULATION

Let  $\mathbf{X}$  denote a  $l_1 \times \dots \times l_n$  unknown  $n$ -dimensional pixelated image to be recovered (e.g.  $n = 2$  or  $n = 3$ ). This image is available as [32] collection of  $P$  projections  $\mathbf{y} = [y_1, \dots, y_P]^T$  which [34] [33] follows the model:

$$\mathbf{y} = T(\boldsymbol{\kappa}, \mathbf{X}) + \mathbf{n}, \quad (1)$$

where  $T(\cdot, \cdot)$  stands for a bilinear function,  $\mathbf{n}$  is a  $P \times 1$  dimension noise vector and  $\boldsymbol{\kappa}$  is the kernel that characterizes the response of the imaging device. [35] Typical point spread responses  $\boldsymbol{\kappa}$  of MRFM tip can be found in [39] for horizontal and vertical configurations. In (1),  $\mathbf{n}$  is an additive Gaussian noise sequence distributed according to  $\mathbf{n} \sim \mathcal{N}(\mathbf{0}, \sigma^2 \mathbf{I}_P)$ .

Note that in standard deblurring problems, the function  $T(\cdot, \cdot)$  represents the standard  $n$ -dimensional convolution operator  $\otimes$ . In this case, the image  $\mathbf{X}$  can be vectorized yielding the unknown image  $\mathbf{x} \in \mathbb{R}^M$  with  $M = P = l_1 l_2 \dots l_n$ . With this notation, Eq. (1) can be rewritten:

$$\mathbf{y} = \mathbf{H}\mathbf{x} + \mathbf{n} \quad \text{or} \quad \mathbf{Y} = \boldsymbol{\kappa} \otimes \mathbf{X} + \mathbf{N} \quad (2)$$

where  $\mathbf{y}$  (resp.  $\mathbf{n}$ ) stands for the vectorized version of  $\mathbf{Y}$  (resp.  $\mathbf{N}$ ) and  $\mathbf{H}$  is an  $P \times M$  matrix that describes convolution by the psf  $\boldsymbol{\kappa}$ .

The problem addressed in the following sections consists of estimating  $\mathbf{x}$  under sparsity and positivity constraints on  $\mathbf{x}$  given the observations  $\mathbf{y}$ , the psf  $\boldsymbol{\kappa}$  and the bilinear function<sup>2</sup>  $T(\cdot, \cdot)$ .

## III. HIERARCHICAL BAYESIAN MODEL

### A. Likelihood function

The observation model defined in (1) and the Gaussian properties of the noise sequence  $\mathbf{n}$  yield:

$$f(\mathbf{y}|\mathbf{x}, \sigma^2) = \left( \frac{1}{2\pi\sigma^2} \right)^P \exp \left( -\frac{\|\mathbf{y} - T(\boldsymbol{\kappa}, \mathbf{x})\|^2}{2\sigma^2} \right), \quad (3)$$

where  $\|\cdot\|$  denotes the standard  $\ell_2$  norm:  $\|\mathbf{x}\|^2 = \mathbf{x}^T \mathbf{x}$ .

### B. Parameter prior distributions

The unknown parameter vector associated with the observation model defined in (1) is  $\boldsymbol{\theta} = \{\mathbf{x}, \sigma^2\}$ . In this section, we introduce prior distributions for these two parameters; which are assumed to be independent.

1) *Image prior*: First let consider the exponential distribution with shape parameter  $a > 0$ :

$$g_a(x_i) = \frac{1}{a} \exp\left(-\frac{x_i}{a}\right) \mathbf{1}_{\mathbb{R}_+^*}(x_i), \quad (4)$$

where  $\mathbf{1}_{\mathbb{E}}(x)$  is the indicator function defined on  $\mathbb{E}$ :

$$\mathbf{1}_{\mathbb{E}}(x) = \begin{cases} 1, & \text{if } x \in \mathbb{E}, \\ 0, & \text{otherwise.} \end{cases} \quad (5)$$

Choosing  $g_a(\cdot)$  as prior distributions for  $x_i$  ( $i = 1, \dots, M$ ) leads to a MAP estimator of  $\mathbf{x}$  that corresponds to a maximum  $\ell_1$ -penalized likelihood estimate with a positivity constraint<sup>3</sup>. Indeed, assuming the component  $x_i$  ( $i = 1, \dots, P$ ) a priori independent allows one to write the full prior distribution for  $\mathbf{x} = [x_1, \dots, x_M]^T$ :

$$g_a(\mathbf{x}) = \left( \frac{1}{a} \right)^M \exp\left(-\frac{\|\mathbf{x}\|_1}{a}\right) \mathbf{1}_{\{\mathbf{x} > 0\}}(\mathbf{x}), \quad (6)$$

where  $\{\mathbf{x} > 0\} = \{\mathbf{x} \in \mathbb{R}^M; x_i > 0, \forall i = 1, \dots, M\}$  and  $\|\cdot\|_1$  is the standard  $\ell_1$  norm  $\|\mathbf{x}\|_1 = \sum_i |x_i|$ . This estimator has shown interesting sparse properties for Bayesian estimation [41] and signal representation [42].

Coupling a standard probability density function (pdf) with an atom at zero is another [36] classical alternative to ensure sparsity. This strategy has for instance been used for located event detection [14] such as spike train deconvolution [11], [17]. In order to increase the sparsity of the prior, we propose

<sup>2</sup>In the following, for sake of conciseness, the same notation  $T(\cdot, \cdot)$  will be adopted for the bilinear operations used on  $n$ -dimensional images  $\mathbf{X}$  and used on  $M \times 1$  vectorized images  $\mathbf{x}$ .

<sup>3</sup>Note that a similar estimator using a Laplacian prior for  $x_i$  ( $i = 1, \dots, M$ ) was proposed in [40] for regression problems and is usually referred to as the LASSO estimator but without positivity constraint.

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extensive performance comparison

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In Section VI we apply our hierarchical Bayesian method to reconstruction of a tobacco virus from real MRFM data.

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to use the following distribution derived from  $g_a(\cdot)$  as prior distribution for  $x_i$ :

$$f(x_i|w, a) = (1-w)\delta(x_i) + wg_a(x_i), \quad (7)$$

where  $\delta(\cdot)$  is the Dirac function. This prior is similar to the LAZE distribution (Laplacian pdf and an atom at zero) introduced in [21] and used, for example, in [22], [23] for MRFM. However, the proposed prior in (7) allows one to take into account the positivity of the pixel value to be estimated. By assuming the components  $x_i$  to be a priori independent ( $i = 1, \dots, M$ ), the following prior distribution is obtained for  $\mathbf{x}$ :

$$f(\mathbf{x}|w, a) = \prod_{i=1}^M [(1-w)\delta(x_i) + wg_a(x_i)]. \quad (8)$$

Introducing the index subsets  $\mathcal{I}_0 = \{i; x_i = 0\}$  and  $\mathcal{I}_1 = \overline{\mathcal{I}_0} = \{i; x_i \neq 0\}$  allows one to rewrite the previous equation as follows:

$$f(\mathbf{x}|w, a) = \left[ (1-w)^{n_0} \prod_{i \in \mathcal{I}_0} \delta(x_i) \right] \left[ w^{n_1} \prod_{i \in \mathcal{I}_1} g_a(x_i) \right], \quad (9)$$

with  $n_\epsilon = \text{card}\{\mathcal{I}_\epsilon\}$ ,  $\epsilon \in \{0, 1\}$ . Note that  $n_0 = M - n_1$  and  $n_1 = \|\mathbf{x}\|_0$  where  $\|\cdot\|_0$  is the standard  $\ell_0$  norm  $\|\mathbf{x}\|_0 = \#\{i; x_i \neq 0\}$ .

2) *Noise variance prior*: A conjugate inverse-Gamma distribution with parameters  $\frac{\nu}{2}$  and  $\frac{\gamma}{2}$  is chosen as prior distribution for the noise variance [43, Appendix A]:

$$\sigma^2|\nu, \gamma \sim \text{IG}\left(\frac{\nu}{2}, \frac{\gamma}{2}\right). \quad (10)$$

In the following,  $\nu$  will be fixed to  $\nu = 2$  and  $\gamma$  will be an hyperparameter to be estimated (see [28], [30], [44]).

### C. Hyperparameter priors

The hyperparameter vector associated with the previous prior distributions is  $\Phi = \{a, \gamma, w\}$ . Obviously, the accuracy of the proposed Bayesian model depends on the values of these hyperparameters. If prior knowledge, e.g. mean number of the non-zero pixels, is available, these parameters can be tuned manually to their actual values. However, in practical situations, such prior information is not available. In this case, as outlined in Section I, these hyperparameters can be estimated directly from the data. Priors for these hyperparameters, sometimes referred to as ‘‘hyperpriors’’ are detailed below.

1) *Hyperparameter  $a$* : A conjugate inverse-Gamma distribution is assumed for hyperparameter  $a$ :

$$a|\alpha \sim \text{IG}(\alpha_0, \alpha_1), \quad (11)$$

with  $\alpha = [\alpha_0, \alpha_1]^T$ . The fixed hyperparameters  $\alpha_0$  and  $\alpha_1$  have been chosen to obtain a vague prior:  $\alpha_0 = \alpha_1 = 10^{-10}$  (see for example [45] for a similar choice).

2) *Hyperparameter  $\gamma$* : A [24] informative Jeffreys’ prior [46], [47] is assumed for hyperparameter  $\gamma$ :

$$f(\gamma) \propto \frac{1}{\gamma} \mathbf{1}_{\mathbb{R}^+}(\gamma). \quad (12)$$

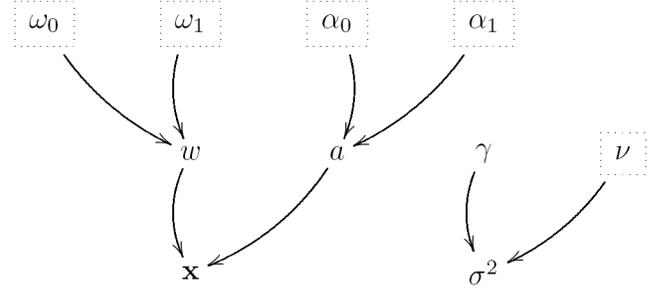


Fig. 1. DAG for the parameter priors and hyperpriors (the fixed hyperparameters appear in dashed boxes).

3) *Hyperparameter  $w$* : A conjugate beta distribution with fixed hyperparameters  $\omega_1$  and  $\omega_0$  is chosen as prior distribution for  $w$ :

$$w|\omega \sim \mathcal{B}(\omega_1, \omega_0), \quad (13)$$

with  $\omega = [\omega_0, \omega_1]^T$  and where  $\mathcal{B}(a, b)$  denotes the Beta distribution with parameters  $(a, b)$ . Note that by choosing  $\omega_0 = \omega_1 = 1$ , the Beta distribution reduces to the uniform distribution on  $[0, 1]$ , which gives the least informative prior.

Assuming that the individual hyperparameters are independent the full hyperparameter prior distribution for  $\Phi$  can be expressed as:

$$\begin{aligned} f(\Phi|\alpha, \omega) &= f(w) f(\gamma) f(a) \\ &= \frac{w^{\omega_1-1} (1-w)^{\omega_0-1}}{aw\gamma B(\omega_1, \omega_0)} \mathbf{1}_{[0,1]}(w) \mathbf{1}_{\mathbb{R}^+}(a) \mathbf{1}_{\mathbb{R}^+}(\gamma), \end{aligned} \quad (14)$$

with  $B(\omega_1, \omega_0) = \frac{\Gamma(\omega_1)\Gamma(\omega_0)}{\Gamma(\omega_1+\omega_0)}$ , where  $\Gamma(\cdot)$  denotes the Gamma function.

### D. Posterior distribution

The posterior distribution of  $\{\theta, \Phi\}$  can be computed as follows:

$$f(\theta, \Phi|\mathbf{y}, \alpha, \omega) \propto f(\mathbf{y}|\theta) f(\theta|\Phi) f(\Phi|\alpha, \omega), \quad (15)$$

with

$$f(\theta|\Phi) = f(\mathbf{x}|a, w) f(\sigma^2|\gamma), \quad (16)$$

where  $f(\mathbf{y}|\theta)$  and  $f(\Phi|\alpha, \omega)$  have been defined in (3) and (14). This hierarchical structure, represented on the directed acyclic graph (DAG) in Fig. 1, allows one to integrate out the parameter  $\sigma^2$  and the hyperparameter vector  $\Phi$  in the full posterior distribution (15), yielding:

$$f(\mathbf{x}|\mathbf{y}, \alpha, \omega) \propto \frac{B(\omega_1 + n_1, \omega_0 + n_0)}{\|\mathbf{y} - T(\kappa, \mathbf{x})\|^P} \frac{\Gamma(n_1 + \alpha_0)}{[\|\mathbf{x}\|_1 + \alpha_1]^{n_1 + \alpha_0}}. \quad (17)$$

where, as defined in paragraph III-B1,  $n_1 = \|\mathbf{x}\|_0$  and  $n_0 = M - \|\mathbf{x}\|_0$ .

The next section presents an appropriate Gibbs sampling strategy [24] that allows one to generate samples distributed according to the posterior distribution  $f(\mathbf{x}|\mathbf{y}, \alpha, \omega)$ .

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#### IV. A GIBBS SAMPLING STRATEGY FOR SPARSE IMAGE RECONSTRUCTION

We propose in this section a Gibbs sampling strategy that allows one to generate samples  $\{\mathbf{x}^{(t)}\}_{t=1,\dots}$  distributed according to the posterior distribution in (17). As simulating directly according to (17) is a difficult task, it is much more convenient to generate samples distributed according to the joint posterior  $f(\mathbf{x}, \sigma^2 | \mathbf{y}, \boldsymbol{\alpha}, \boldsymbol{\omega})$ . The main steps of this algorithm are detailed in subsections IV-A and IV-B (see also Algorithm 1 below).

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##### ALGORITHM 1:

###### Gibbs sampling algorithm for sparse image reconstruction

- **Initialization:**
    - Sample parameter  $\mathbf{x}^{(0)}$  from pdf in (9),
    - Sample parameters  $\tilde{\sigma}^{2(0)}$  from the pdf in (10),
    - Set  $t \leftarrow 1$ ,
  - **Iterations:** for  $t = 1, 2, \dots$ , do
    1. Sample hyperparameter  $w^{(t)}$  from the pdf in (19),
    2. Sample hyperparameter  $a^{(t)}$  from the pdf in (20),
    3. For  $i = 1, \dots, M$ , sample parameter  $x_i^{(t)}$  from pdf in (21),
    4. Sample parameter  $\tilde{\sigma}^{2(t)}$  from the pdf in (24),
    5. Set  $t \leftarrow t + 1$ .
- 
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##### A. Generation of samples according to $f(\mathbf{x} | \sigma^2, \mathbf{y}, \boldsymbol{\alpha}, \boldsymbol{\omega})$

To generate samples distributed according to  $f(\mathbf{x} | \sigma^2, \mathbf{y}, \boldsymbol{\omega})$ , it is very convenient to sample according to  $f(\mathbf{x}, w, a | \sigma^2, \mathbf{y}, \boldsymbol{\omega})$  in the following 3-step procedure.

1) *Generation of samples according to  $f(w | \mathbf{x}, \boldsymbol{\omega})$ :* Using (9), the following result can be obtained:

$$f(w | \mathbf{x}, \boldsymbol{\omega}) \propto (1 - w)^{n_0 + \omega_0 - 1} w^{n_1 + \omega_1 - 1}, \quad (18)$$

where  $n_0$  and  $n_1$  have been defined in paragraph III-B1. Therefore, generation of samples according to  $f(w | \mathbf{x}, \boldsymbol{\omega})$  is achieved as follows:

$$w | \mathbf{x}, \boldsymbol{\omega} \sim \mathcal{B}e(\omega_1 + n_1, \omega_0 + n_0). \quad (19)$$

2) *Generation of samples according to  $f(a | \mathbf{x}, \boldsymbol{\alpha})$ :* Looking at the joint posterior distribution (15), it yields:

$$a | \mathbf{x}, \boldsymbol{\alpha} \sim \mathcal{IG}(\|\mathbf{x}\|_0 + \alpha_0, \|\mathbf{x}\|_1 + \alpha_1). \quad (20)$$

3) *Generation of samples according to  $f(\mathbf{x} | w, a, \sigma^2, \mathbf{y})$ :* The prior chosen for  $x_i$  ( $i = 1, \dots, M$ ) yields a posterior distribution of  $\mathbf{x}$  that is not closed form. However, the posterior distribution of each component  $x_i$  ( $i = 1, \dots, M$ ) conditionally upon the others can be easily derived. Indeed straightforward computations detailed in Appendix A yield:

$$f(x_i | w, a, \sigma^2, \mathbf{x}_{-i}, \mathbf{y}) \propto (1 - w_i) \delta(x_i) + w_i \phi_+(x_i | \mu_i, \eta_i^2), \quad (21)$$

where  $\mathbf{x}_{-i}$  stands for the vector  $\mathbf{x}$  whose  $i$ th component has been removed and  $\mu_i$  and  $\eta_i^2$  are given in Appendix A. In

(21),  $\phi_+(\cdot, m, s^2)$  stands for the pdf of the truncated Gaussian distribution defined on  $\mathbb{R}_+^*$  with hidden parameters equal to mean  $m$  and variance  $s^2$ :

$$\phi_+(x, m, s^2) = \frac{1}{C(m, s^2)} \exp\left[-\frac{(x - m)^2}{2s^2}\right] \mathbf{1}_{\mathbb{R}_+^*}(x), \quad (22)$$

with

$$C(m, s^2) = \sqrt{\frac{\pi s^2}{2}} \left[1 + \operatorname{erf}\left(\frac{m}{\sqrt{2s^2}}\right)\right]. \quad (23)$$

The form in (21) specifies  $x_i | w, a, \sigma^2, \mathbf{x}_{-i}, \mathbf{y}$  as a Bernoulli-truncated Gaussian variable with parameter  $(w_i, \mu_i, \eta_i^2)$ . Appendix C presents an algorithm that can be used to generate samples distributed according to this distribution.

To summarize, generating samples distributed according to  $f(\mathbf{x} | w, \sigma^2, a, \mathbf{y})$  can be performed by updating the coordinates of  $\mathbf{x}$  successively using  $M$  Gibbs moves (requiring generation of Bernoulli-truncated Gaussian variables).

##### B. Generation of samples according to $f(\sigma^2 | \mathbf{x}, \mathbf{y})$

Samples are generated as the following way:

$$\sigma^2 | \mathbf{x}, \mathbf{y} \sim \mathcal{IG}\left(\frac{P}{2}, \frac{\|\mathbf{y} - T(\boldsymbol{\kappa}, \mathbf{x})\|^2}{2}\right). \quad (24)$$

#### V. SIMULATION ON SYNTHETIC IMAGES

TABLE I  
PARAMETERS USED TO COMPUTE THE MRFM PSF.

Parameter		Value
Description	Name	
Amplitude of external magnetic field	$B_{\text{ext}}$	$9.4 \times 10^3$ G
Value of $B_{\text{mag}}$ in the resonant slice	$B_{\text{res}}$	$1.0 \times 10^4$ G
Radius of tip	$R_0$	4.0 nm
Distance from tip to sample	$d$	6.0 nm
Cantilever tip moment	$m$	$4.6 \times 10^5$ emu
Peak cantilever oscillation	$x_{\text{pk}}$	0.8 nm
Maximum magnetic field gradient	$G_{\text{max}}$	125

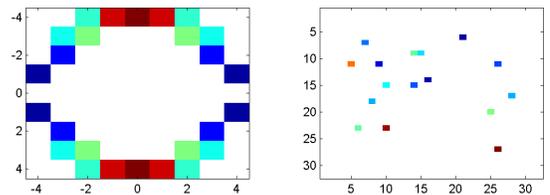


Fig. 2. Left: Psf of the MRFM tip. Right: unknown sparse image to be estimated.

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### A. Reconstruction of 2-dimensional image

In this subsection, a  $32 \times 32$  synthetic image, depicted in Fig. 2 (right), is simulated using the prior in (9) with parameter  $a = 1$  and  $w = 0.02$ . In this figure and in the following ones, white pixels stands for identically null values. A general analytical derivation of the psf of the MRFM tip has been given in [39] and is explained in [23]. Following this model, a  $10 \times 10$  2-dimensional convolution kernel, represented in Fig. 2 (left), has been generated when the physical parameters are tuned to the values gathered in Table I. The corresponding matrix  $\mathbf{H}$  introduced in (2) is of size  $1024 \times 1024$ . The observed measurements  $\mathbf{y}$ , depicted in Fig. 2 (right) are of size  $P = 1024$ . These observations are corrupted by an additive Gaussian noise with two different variances  $\sigma^2 = 1.2 \times 10^{-1}$  and  $\sigma^2 = 1.6 \times 10^{-3}$ , corresponding to signal-to-noise ratios  $\text{SNR} = 2\text{dB}$  and  $\text{SNR} = 20\text{dB}$  respectively.

1) *Simulation results:* The observations are processed by the proposed algorithm that consists of  $N_{\text{MC}} = 2000$  iterations of the Gibbs sampler with  $N_{\text{bi}} = 300$  burn-in iterations. Then the MAP estimator of the unknown image  $\mathbf{x}$  is computed by keeping among  $\mathcal{X} = \{\mathbf{x}^{(t)}\}_{t=1, \dots, N_{\text{MC}}}$  the generated sample that maximizes the posterior distribution in (17):

$$\begin{aligned} \hat{\mathbf{x}}_{\text{MAP}} &= \underset{\mathbf{x} \in \mathbb{R}_+^M}{\text{argmax}} f(\mathbf{x}|\mathbf{y}) \\ &\approx \underset{\mathbf{x} \in \mathcal{X}}{\text{argmax}} f(\mathbf{x}|\mathbf{y}). \end{aligned} \quad (25)$$

These estimates are depicted in Fig. 3 for the two levels of noise considered. It can be noticed that the estimated image is very similar to the actual image, even with a low SNR.

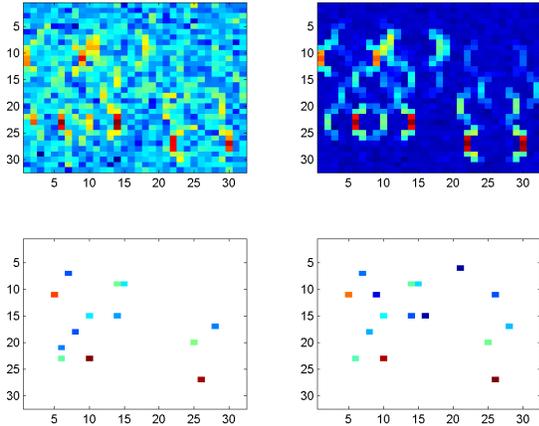


Fig. 3. Top, left (resp. right): noisy observations for  $\text{SNR} = 2\text{dB}$  (resp.  $20\text{dB}$ ). Bottom, left (resp. right): reconstructed image for  $\text{SNR} = 2\text{dB}$  (resp.  $20\text{dB}$ ).

Moreover, as the proposed algorithm generates samples distributed according to the posterior distribution in (17), these samples can be used to compute the posterior distributions of each parameter. As examples, the posterior distributions of the hyperparameters  $a$  and  $w$ , as well as the noise variance  $\sigma^2$ , are shown in Fig. 4, 5 and 6. These estimated distributions are in good agreement with the actual values of these parameters for the two SNR levels.

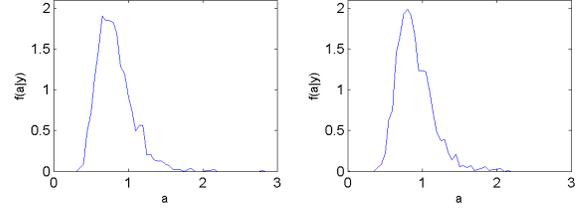


Fig. 4. Posterior distribution of hyperparameter  $a$  (left:  $\text{SNR} = 2\text{dB}$ , right:  $\text{SNR} = 20\text{dB}$ ).

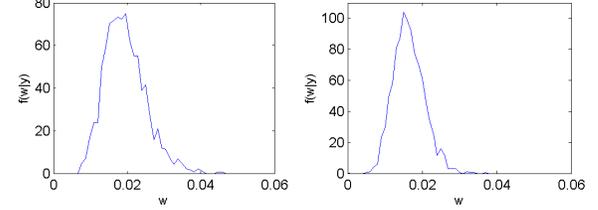


Fig. 5. Posterior distribution of hyperparameter  $w$  (left:  $\text{SNR} = 2\text{dB}$ , right:  $\text{SNR} = 20\text{dB}$ ).

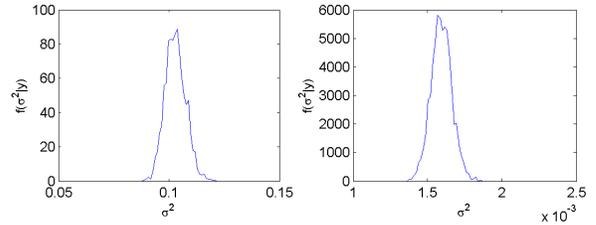


Fig. 6. Posterior distribution of hyperparameter  $\sigma^2$  (left:  $\text{SNR} = 2\text{dB}$ , right:  $\text{SNR} = 20\text{dB}$ ).

The posterior distributions of four different pixels are depicted in Fig. 7. These posteriors are also in agreement with the actual values of these pixels that are represented in dotted red line in these figures.

2) *Comparison of reconstruction performances:* The results provided by the proposed method have been compared with those provided by methods that also estimate the hyperparameters automatically. Firstly, the techniques proposed in [22], [23] are based on EM algorithms that perform empirical estimation of the unknown hyperparameters. Therein, two empirical Bayesian estimators, denoted Emp-MAP-Lap and Emp-MAP-LAZE, based on a Laplacian or a LAZE prior respectively, are studied. Here we compare the estimators of [22], [23] to the MMSE estimator and the MAP estimator under the model and the algorithm presented in Sections III and IV. The MMSE estimator of the unknown parameter  $\mathbf{x}$  is obtained by empirical averaging over the last  $N_r = 1700$  outputs of the sampler according to:

$$\begin{aligned} \hat{\mathbf{x}}_{\text{MMSE}} &= \text{E}[\mathbf{x}|\mathbf{y}] \\ &\approx \frac{1}{N_r} \sum_{t=1}^{N_r} \mathbf{x}^{(N_{\text{bi}}+t)}. \end{aligned} \quad (26)$$

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	Number: 7 Author: vmuser Here we compare our proposed hierarchical Bayesian method to the methods of [22],[23].	Subject: Replacement Text	Date: 9/17/2008 1:00:42 PM
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	Number: 13 Author: vmuser	Subject: Cross-Out	Date: 9/17/2008 1:02:25 PM
	Number: 14 Author: vmuser Gibbs	Subject: Inserted Text	Date: 9/17/2008 1:02:28 PM
	Number: 15 Author: vmuser ground truth	Subject: Replacement Text	Date: 9/17/2008 12:59:21 PM
	Number: 16 Author: vmuser	Subject: Cross-Out	Date: 9/17/2008 12:59:31 PM
	Number: 17 Author: vmuser ; randomly drawn from the prior distribution	Subject: Inserted Text	Date: 9/17/2008 12:59:30 PM

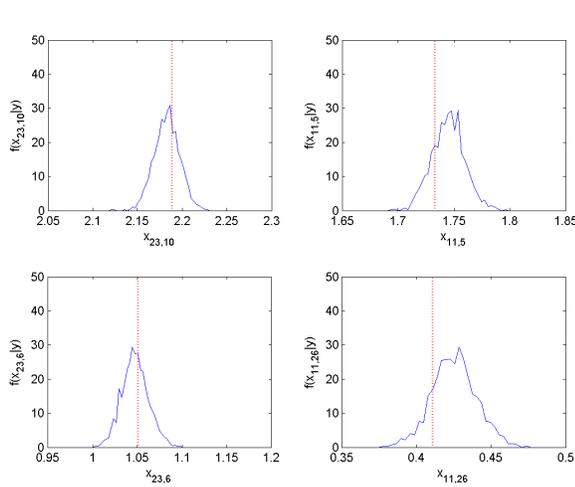


Fig. 7. Posteriors distributions of the non-zero values of  $\mathbf{x}$  for SNR = 20dB, (actual values are depicted with dotted red lines).

Finally, the results are compared with the estimator provided by a standard Landweber algorithm [48]. The proposed comparison is conducted with respect to several measures of performance. First let  $\mathbf{e} = \mathbf{x} - \hat{\mathbf{x}}$  denote the reconstruction error when  $\hat{\mathbf{x}}$  is the estimator of the image  $\mathbf{x}$  to be recovered. To measure the performance of the sparse reconstruction, criteria inspired by [23] have been used: the  $\ell_0$ ,  $\ell_1$  and  $\ell_2$ -norms of  $\mathbf{e}$  measure accuracy of the reconstruction and the  $\ell_0$ -norm of the estimator  $\hat{\mathbf{x}}$  measure its sparsity. Moreover, as noticed in [23], small non-zero values of the pixel are usually not distinguishable from exactly zero values by a human being. Following this remark, a less strict measure of sparsity than the  $\ell_0$ -norm has been introduced. This measure<sup>4</sup>, denoted  $\|\cdot\|_\delta$ , is the number of components that are less than a given threshold  $\delta$ :

$$\begin{aligned} \|\hat{\mathbf{x}}\|_\delta &= \sum_{i=1}^M \mathbf{1}_{\hat{x}_i < \delta}(\hat{x}_i), \\ \|\mathbf{e}\|_\delta &= \sum_{i=1}^M \mathbf{1}_{e_i < \delta}(e_i). \end{aligned} \quad (27)$$

It what follows,  $\delta$  has been chosen as  $\delta = 10^{-2} \|\mathbf{x}\|_\infty$ . To summarize, the following criteria have been computed for the image in paragraph V-A1 for two levels of SNR:  $\|\mathbf{e}\|_0$ ,  $\|\mathbf{e}\|_\delta$ ,  $\|\mathbf{e}\|_1$ ,  $\|\mathbf{e}\|_2$ ,  $\|\hat{\mathbf{x}}\|_0$  and  $\|\hat{\mathbf{x}}\|_\delta$ .

Table II gathers the six performance measures for the five different studied algorithms. It clearly appears that the proposed Bayesian method outperforms the others in the  $\ell_1$  or  $\ell_2$ -norm evaluations of the error reconstruction, whatever the estimator chosen (MAP or MMSE). This can be easily explained by the accurate estimation of the hyperparameters thanks to the introduced hierarchical model. It also appears that an MMSE estimation of the unknown image yields to a non sparse estimator in a  $\ell_0$ -norm sense. This can be explained by a very weak posterior probability of having non-zero value

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each pixel, a less decision, by using the sparsity measure  $\|\cdot\|_\delta$ . In this instance, the MAP estimator seems to be a very powerful estimator for the sparse reconstruction problem as it seems to balance the sparsity of the solution and the minimization of the reconstruction error. However, it has to be noticed that MMSE estimation obtains more information than a point estimation and can be useful to derive confidence intervals.

TABLE II  
RECONSTRUCTION PERFORMANCES FOR DIFFERENT SPARSE  
DECONVOLUTION ALGORITHMS.

Method	Error criterion					
	$\ \mathbf{e}\ _0$	$\ \mathbf{e}\ _\delta$	$\ \mathbf{e}\ _1$	$\ \mathbf{e}\ _2$	$\ \hat{\mathbf{x}}\ _0$	$\ \hat{\mathbf{x}}\ _\delta$
SNR = 2dB						
Landweber	1024	990	339.76	13.32	1024	990
Emp-MAP-Lap	18	17	14.13	4.40	0	0
Emp-MAP-LAZE	60	58	9.49	1.44	55	55
Proposed MMSE	1001	30	3.84	0.72	1001	27
Proposed MAP	19	16	2.38	0.81	13	13
SNR = 20dB						
Landweber	1024	931	168.85	6.67	1024	931
Emp-MAP-Lap	33	18	1.27	0.31	28	23
Emp-MAP-LAZE	144	19	1.68	0.22	144	27
Proposed MMSE	541	5	0.36	0.11	541	16
Proposed MAP	19	7	0.39	0.13	16	16

### B. Reconstruction of undersampled 3-dimensional images

In this subsection, some simulation results are presented to illustrate the performance of the algorithm when applied on undersampled 3D images. First, a  $24 \times 24 \times 6$  image is generated such as 4 pixels have non-zero values in each  $z$  slice. The resulting data is depicted in Fig. 8 (right) and Fig. 10 (top). This image to be recovered is assumed to be convolved with a  $5 \times 5 \times 3$  kernel that is represented in Fig. 8 (right). The resulting convolved image is depicted in Fig. 9 (left). However, the actually observed image is assumed to be an undersampled version of this image. More precisely, the sampling rates are assumed to be  $d_x = 2$ ,  $d_y = 3$   $d_z = 1$  respectively in the 3 dimensions. Consequently the observed 3D image, shown in Fig. 9, is of size  $12 \times 8 \times 6$ . Finally, an i.i.d. Gaussian noise with  $\sigma = 0.02$  is added following the model in (1). Note that under these assumptions, the application  $T(\cdot, \cdot)$  can be split into two standard operations following the composition:

$$T(\boldsymbol{\kappa}, \mathbf{X}) = g_{d_x, d_y, d_z}(\boldsymbol{\kappa} \otimes \mathbf{X}), \quad (28)$$

where  $g_{d_x, d_y, d_z}(\cdot)$  stands for the undersampling function.

The proposed Bayesian algorithm is used to perform the sparse reconstruction. The number of Monte Carlo runs has been fixed to  $N_{MC} = 2000$  with  $N_{bi} = 200$  burn-in iterations. The MAP estimator has been chosen as the reconstructed image estimate since it outperforms the MMSE ones (as explained in paragraph V-A2). This MAP estimator, depicted in Fig. 10 (bottom), is quite satisfactory given the problem difficulty introduced by the undersampling.

	Number: 1 Author: vmuser	Subject: Cross-Out	Date: 9/17/2008 1:06:34 PM
	Number: 2 Author: vmuser at many	Subject: Replacement Text	Date: 9/17/2008 1:06:25 PM
	Number: 3 Author: vmuser s	Subject: Inserted Text	Date: 9/17/2008 1:06:25 PM
	Number: 4 Author: vmuser The	Subject: Inserted Text	Date: 9/17/2008 1:06:38 PM
	Number: 5 Author: vmuser	Subject: Cross-Out	Date: 9/17/2008 1:06:42 PM
	Number: 6 Author: vmuser indicates that most of the pixels are in fact very close to zero.	Subject: Inserted Text	Date: 9/17/2008 1:06:53 PM
	Number: 7 Author: vmuser	Subject: Cross-Out	Date: 9/17/2008 1:07:06 PM
	Number: 8 Author: vmuser The	Subject: Replacement Text	Date: 9/17/2008 1:07:02 PM
	Number: 9 Author: vmuser by construction the	Subject: Replacement Text	Date: 9/17/2008 1:07:28 PM
	Number: 10 Author: vmuser will always have lower mean square error	Subject: Replacement Text	Date: 9/17/2008 1:07:42 PM
	Number: 11 Author: vmuser	Subject: Pencil	Date: 9/17/2008 1:02:56 PM
	Number: 12 Author: vmuser We also compare	Subject: Replacement Text	Date: 9/17/2008 1:02:50 PM
	Number: 13 Author: vmuser	Subject: Cross-Out	Date: 9/17/2008 1:03:24 PM
	Number: 14 Author: vmuser As in [23] we compare estimators	Subject: Inserted Text	Date: 9/17/2008 1:03:33 PM
	Number: 15 Author: vmuser	Subject: Cross-Out	Date: 9/17/2008 1:03:24 PM
	Number: 16 Author: vmuser	Subject: Cross-Out	Date: 9/17/2008 1:03:34 PM
	Number: 17 Author: vmuser	Subject: Cross-Out	Date: 9/17/2008 1:03:40 PM
	Number: 18 Author: vmuser criteria	Subject: Inserted Text	Date: 9/17/2008 1:03:38 PM
	Number: 19 Author: vmuser Let	Subject: Inserted Text	Date: 9/17/2008 1:03:42 PM
	Number: 20 Author: vmuser These criteria are:	Subject: Replacement Text	Date: 9/17/2008 1:03:53 PM
	Number: 21 Author: vmuser	Subject: Cross-Out	Date: 9/17/2008 1:03:56 PM
	Number: 22 Author: vmuser , which	Subject: Inserted Text	Date: 9/17/2008 1:03:55 PM
	Number: 23 Author: vmuser s	Subject: Inserted Text	Date: 9/17/2008 1:04:06 PM
	Number: 24 Author: vmuser	Subject: Cross-Out	Date: 9/17/2008 1:04:01 PM
	Number: 25 Author: vmuser	Subject: Cross-Out	Date: 9/17/2008 1:04:00 PM
	Number: 26 Author: vmuser , which	Subject: Inserted Text	Date: 9/17/2008 1:04:04 PM
	Number: 27 Author: vmuser ,	Subject: Inserted Text	Date: 9/17/2008 1:03:57 PM
	Number: 28 Author: vmuser s	Subject: Inserted Text	Date: 9/17/2008 1:04:04 PM

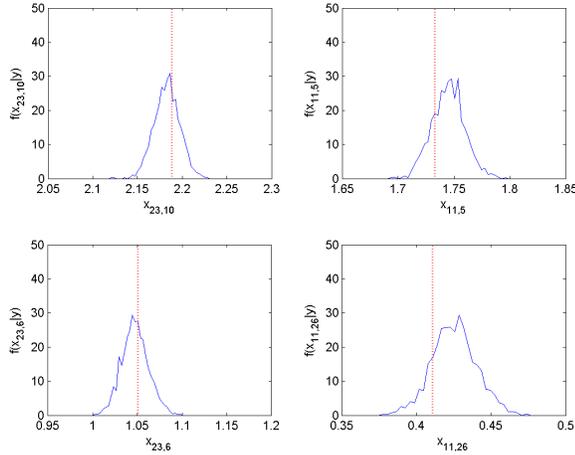


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	Number: 29 Author: vmuser	Subject: Replacement Text	Date: 9/17/2008 1:04:14 PM
	As pointed out		
	Number: 30 Author: vmuser	Subject: Replacement Text	Date: 9/17/2008 1:04:10 PM
	Thus		
	Number: 31 Author: vmuser	Subject: Inserted Text	Date: 9/17/2008 1:08:43 PM
	As discussed in Sec. VI, the prototype MRFM instrument collects data projections as irregularly spaced, or undersampled, spatial samples.		
	Number: 32 Author: vmuser	Subject: Replacement Text	Date: 9/17/2008 1:09:11 PM
	we indicate how the image reconstruction algorithm can be adapted to this undersampled scenario in 3D.		
	Number: 33 Author: vmuser	Subject: Cross-Out	Date: 9/17/2008 1:04:25 PM
	Number: 34 Author: vmuser	Subject: Cross-Out	Date: 9/17/2008 1:09:13 PM
	Number: 35 Author: vmuser	Subject: Inserted Text	Date: 9/17/2008 1:04:26 PM
	,		
	Number: 36 Author: vmuser	Subject: Inserted Text	Date: 9/17/2008 1:09:24 PM
	For concreteness, we illustrate by a concrete example.		
	Number: 37 Author: vmuser	Subject: Cross-Out	Date: 9/17/2008 1:09:13 PM
	Number: 38 Author: vmuser	Subject: Replacement Text	Date: 9/17/2008 1:04:33 PM
	which is		
	Number: 39 Author: vmuser	Subject: Replacement Text	Date: 9/17/2008 1:04:43 PM
	reconstructed image pixels		
	Number: 40 Author: vmuser	Subject: Cross-Out	Date: 9/17/2008 1:09:28 PM
	Number: 41 Author: vmuser	Subject: Inserted Text	Date: 9/17/2008 1:09:31 PM
	an		
	Number: 42 Author: vmuser	Subject: Replacement Text	Date: 9/17/2008 1:04:49 PM
	shows		
	Number: 43 Author: vmuser	Subject: Cross-Out	Date: 9/17/2008 1:04:51 PM
	Number: 44 Author: vmuser	Subject: Inserted Text	Date: 9/17/2008 1:04:59 PM
	studied		
	Number: 45 Author: vmuser	Subject: Replacement Text	Date: 9/17/2008 1:05:03 PM
	The		
	Number: 46 Author: vmuser	Subject: Cross-Out	Date: 9/17/2008 1:05:15 PM
	Number: 47 Author: vmuser	Subject: Inserted Text	Date: 9/17/2008 1:05:14 PM
	s (labeled "proposed MMSE" and "proposed MAP" in the table)		
	Number: 48 Author: vmuser	Subject: Replacement Text	Date: 9/17/2008 1:05:19 PM
	estimators		
	Number: 49 Author: vmuser	Subject: Replacement Text	Date: 9/17/2008 1:05:28 PM
	terms of		
	Number: 50 Author: vmuser	Subject: Cross-Out	Date: 9/17/2008 1:05:31 PM
	Number: 51 Author: vmuser	Subject: Inserted Text	Date: 9/17/2008 1:05:35 PM
	s		
	Number: 52 Author: vmuser	Subject: Cross-Out	Date: 9/17/2008 1:09:50 PM
	Number: 53 Author: vmuser	Subject: Replacement Text	Date: 9/17/2008 1:05:50 PM
	Note that the		
	Number: 54 Author: vmuser	Subject: Cross-Out	Date: 9/17/2008 1:10:24 PM
	Number: 55 Author: vmuser	Subject: Cross-Out	Date: 9/17/2008 1:10:24 PM
	Number: 56 Author: vmuser	Subject: Inserted Text	Date: 9/17/2008 1:10:30 PM
	For illustration the proposed		
	Number: 57 Author: vmuser	Subject: Inserted Text	Date: 9/17/2008 1:09:43 PM
	hierarchical		
	Number: 58 Author: vmuser	Subject: Inserted Text	Date: 9/17/2008 1:10:37 PM
	with undersampled data		

Comments from page 6 continued on next page

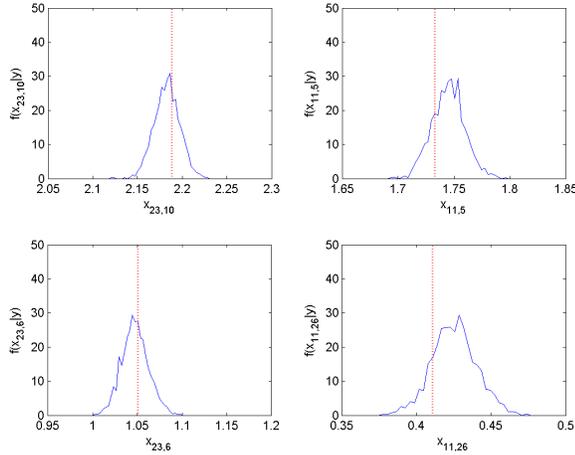


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-  Number: 59 Author: vmuser Subject: Replacement Text Date: 9/17/2008 1:10:40 PM  
was
- 
-  Number: 60 Author: vmuser Subject: Cross-Out Date: 9/17/2008 1:05:44 PM
- 
-  Number: 61 Author: vmuser Subject: Replacement Text Date: 9/17/2008 1:05:55 PM  
is a
- 
-  Number: 62 Author: vmuser Subject: Cross-Out Date: 9/17/2008 1:05:55 PM
- 
-  Number: 63 Author: vmuser Subject: Cross-Out Date: 9/17/2008 1:05:55 PM
- 
-  Number: 64 Author: vmuser Subject: Replacement Text Date: 9/17/2008 1:11:18 PM  
Figure 10 shows the result of applying the proposed MAP estimator to the estimated posterior.
- 
-  Number: 65 Author: vmuser Subject: Inserted Text Date: 9/17/2008 1:05:57 PM  
the
- 
-  Number: 66 Author: vmuser Subject: Replacement Text Date: 9/17/2008 1:06:05 PM  
is due to the
- 
-  Number: 67 Author: vmuser Subject: Cross-Out Date: 9/17/2008 1:06:14 PM
- 
-  Number: 68 Author: vmuser Subject: Replacement Text Date: 9/17/2008 1:06:11 PM  
small
- 
-  Number: 69 Author: vmuser Subject: Inserted Text Date: 9/17/2008 1:06:08 PM  
but non-zero

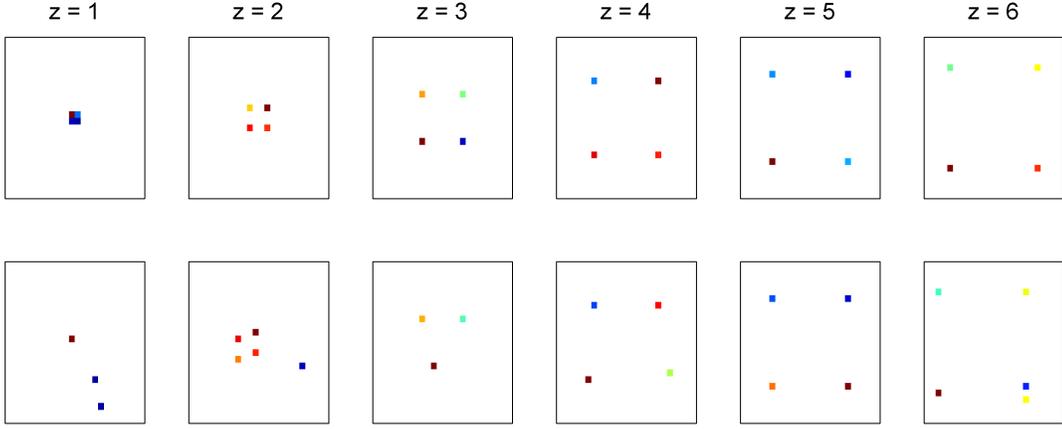


Fig. 10. Top: slices of the sparse image to be recovered. Bottom: slices of the estimated sparse image.

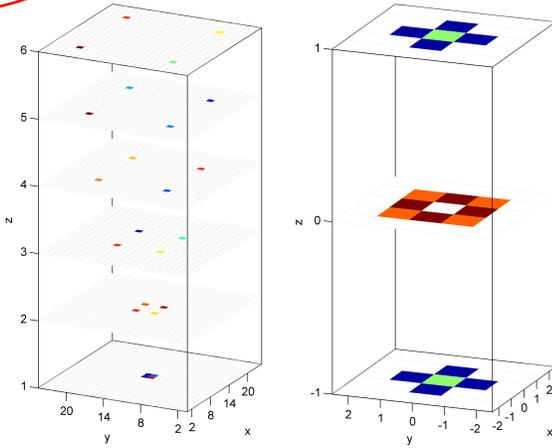


Fig. 8. Left:  $24 \times 24 \times 6$  unknown image to be recovered. Right:  $5 \times 5 \times 3$  kernel modeling the psf.

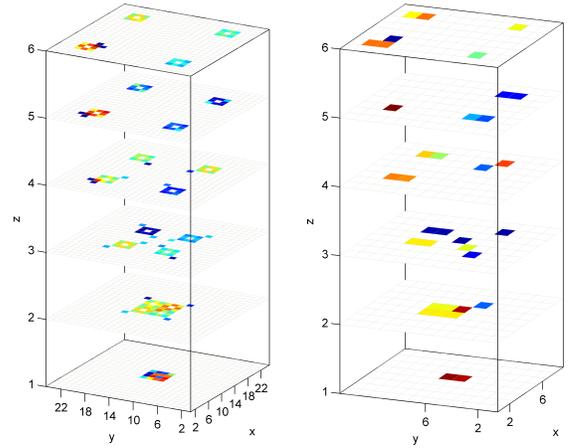


Fig. 9. Left:  $24 \times 24 \times 6$  regularly sampled convolved image. Right:  $12 \times 8 \times 6$  undersampled observed image.

## VI. APPLICATION ON REAL MRFM IMAGES

The 3-dimensional real data used in this section have been initially presented in [35] to illustrate the nanometer spatial resolution of MRFM. The observed sample consists of a collection of Tobacco mosaic virus particles that are divided into a whole segment and other fragments. The signal is computed from the measured proton distribution and the 3-dimensional psf following the protocol described in [35] and [49]. The resulting scan data are depicted in Figure 11 (top) for four different distances between the MRFM tip and the sample:  $d = 24\text{nm}$ ,  $d = 37\text{nm}$ ,  $d = 50\text{nm}$  and  $d = 62\text{nm}$ . Each of these x-y slices is of size  $60 \times 32$ .

These experimental data are undersampled, i.e. the spatial resolution of the MRFM tip, and therefore the psf function, is finer than the resolution of the observed slices. Consequently, these data have been deconvolved taking into account the oversampling rates defined by  $d_x = 3$ ,  $d_y = 2$  and  $d_z = 3$  in the three directions. The MAP estimate of the unknown

image is computed after  $N_{MC} = 1000$  (with  $N_{bi} = 200$ ) the proposed Bayesian algorithm initialized with the output of the Landweber iteration. Three horizontal slices of the estimated image<sup>5</sup> is depicted in Figure 12. A 3-dimensional view of the estimated profile of the virus fragments is also available in Figure 13. The MMSE estimates of the parameters introduced in Section III are  $\hat{\sigma}_{MMSE}^2 = 0.10$ ,  $\hat{a}_{MMSE} = 1.9 \times 10^{-12}$  and  $\hat{w}_{MMSE} = 1.4 \times 10^{-2}$ .

To illustrate the performance of the proposed deconvolution algorithm, the data reconstructed from the estimated 3-dimensional image are depicted in Figure 11 (bottom). These figures are clearly in good agreement with the observed data (top). Moreover, to evaluate the convergence speed, the reconstruction error is represented in Figure 14 as a function of the iterations for the proposed Bayesian and the Landweber algorithms. It clearly appears that the convergence rate of our

<sup>5</sup>Note that most part of the estimated 3 dimensional image is empty space due to the very localized position of the imaged data.

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	Here we illustrate the heirarchical Bayes MAP reconstruction algorithm for real lthree dimensional MRFM data. The data is a set of MRFM projections of a sample of tobacco virus. Comprehensive details of both the experiment and the MRFM data acquisition protocol are given in [35].		
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	Several more iterations of the Landweber algorithm would produce the reconstructions reported in [35]. Theimage reconstructions produced by the Landweber and Bayesian MAP algorithm are shown in Figs. 11-13.		
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	By forward projecting the estimated virus image through the point spread function one can visually evaluate the goodness of fit of the reconstruction to the raw measured data. This is depicted in Fig. 11.		
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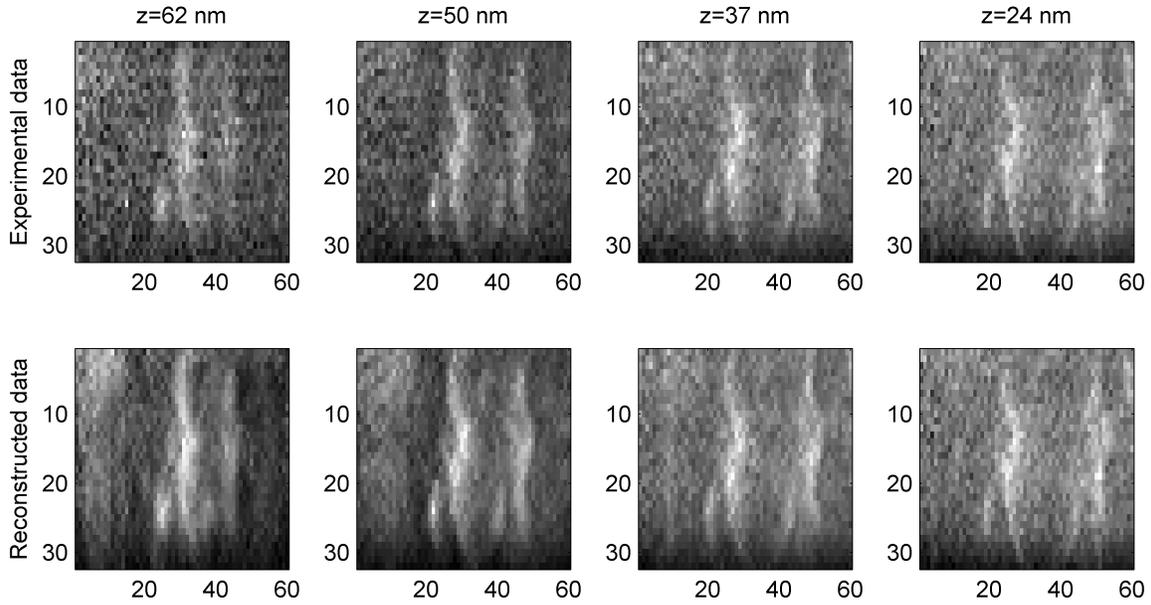


Fig. 11. Top: experimental scan data. Bottom: scan data computed from the proposed Bayesian reconstruction.

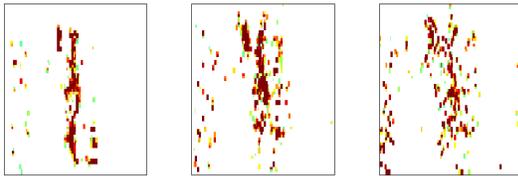


Fig. 12. Three horizontal slices of the estimated image.

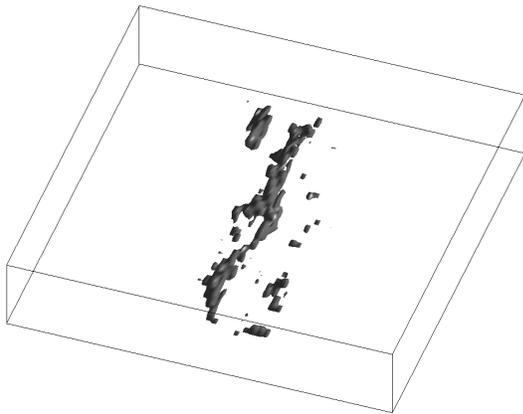


Fig. 13. 3-dimensional view of the estimated profile of the Tobacco virus fragments.

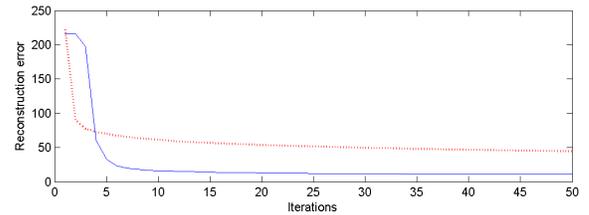


Fig. 14. Error reconstructions as functions of the iteration number for the proposed algorithm (continuous blue line) and Landweber algorithm (dotted red line).

## VII. CONCLUSIONS

This paper presented a Bayesian sampling algorithm for solving deconvolution of sparse images corrupted by additive Gaussian noise. A Bernoulli-truncated exponential distribution was proposed as prior distribution for the sparse image to be recovered. The hyperparameters of the model were estimated in a fully Bayesian scheme by choosing prior distributions for them and by integrating them out from the full posterior distribution. An efficient Gibbs sampler allowed one to generate samples distributed according to this posterior distribution. The derived Bayesian estimators performed significantly better than estimators classically used to solve sparse reconstruction problems. This was mainly due to a performing estimation of the hyperparameters via the proposed hierarchical model.

algorithm is significantly better than the Landweber algorithm.

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	integrated out of the posterior distribution of the image producing a full posterior distribution that can be used for estimation of the pixel values by maximization (MAP) or integration (MMSE).		
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	Our approach was implemented on real MRFM data to form a 3D image of a tobacco virus. Future work will include extension of the proposed method to other sparse bases, inclusion of uncertain point spread functions, and investigation of molecular priors.		

POSTERIOR DISTRIBUTION  $f(x_i | w, a, \sigma^2, \mathbf{x}_{-i}, \mathbf{y})$

The posterior distribution of each component  $x_i$  ( $i = 1, \dots, M$ ) conditionally upon the others is linked to the likelihood function (3) and the prior distribution (7) via the Bayes' paradigm:

$$f(x_i | w, a, \sigma^2, \mathbf{x}_{-i}, \mathbf{y}) \propto f(\mathbf{y} | \mathbf{x}, \sigma^2) f(x_i | w, a). \quad (29)$$

This distribution can be easily derived by decomposing  $\mathbf{x}$  on the standard orthonormal basis

$$\mathbb{B} = \{\mathbf{u}_1, \dots, \mathbf{u}_M\}, \quad (30)$$

where  $\mathbf{u}_i$  is the  $i$ th column of the  $M \times M$  identity matrix. Indeed, let decompose

$$\mathbf{x} = \tilde{\mathbf{x}}_i + x_i \mathbf{u}_i, \quad (31)$$

where  $\tilde{\mathbf{x}}_i$  is the vector  $\mathbf{x}$  whose  $i$ th element has been replaced by 0. Then the linear property of the operator  $T(\boldsymbol{\kappa}, \cdot)$  allows one to state:

$$T(\boldsymbol{\kappa}, \mathbf{x}) = T(\boldsymbol{\kappa}, \tilde{\mathbf{x}}_i) + x_i T(\boldsymbol{\kappa}, \mathbf{u}_i). \quad (32)$$

Consequently, (29) can be rewritten

$$f(x_i | w, a, \sigma^2, \mathbf{x}_{-i}, \mathbf{y}) \propto \exp\left(-\frac{\|\mathbf{e}_i - x_i \mathbf{h}_i\|^2}{2\sigma^2}\right) \times \left[(1-w)\delta(x_i) + \frac{w}{a} \exp\left(-\frac{x_i}{a}\right) \mathbf{1}_{\mathbb{R}_+^*}(x_i)\right], \quad (33)$$

where<sup>6</sup>

$$\begin{cases} \mathbf{e}_i = \mathbf{y} - T(\boldsymbol{\kappa}, \tilde{\mathbf{x}}_i), \\ \mathbf{h}_i = T(\boldsymbol{\kappa}, \mathbf{u}_i). \end{cases} \quad (34)$$

An efficient way to compute  $\mathbf{e}_i$  within the Gibbs sampler scheme is reported in Appendix B. Then, straightforward computations similar to those in [11] and [50, Annex B] yield to the following distribution:

$$f(x_i | w, a, \sigma^2, \mathbf{x}_{-i}, \mathbf{y}) \propto (1-w_i)\delta(x_i) + w_i \phi_+(x_i | \mu_i, \eta_i^2), \quad (35)$$

with

$$\begin{cases} \eta_i^2 = \frac{\sigma^2}{\|\mathbf{h}_i\|^2}, \\ \mu_i = \eta_i^2 \left( \frac{\mathbf{h}_i^T \mathbf{e}_i}{\sigma^2} - \frac{1}{a} \right), \end{cases} \quad (36)$$

and

$$\begin{cases} u_i = \frac{w}{a} C(\mu_i, \eta_i^2) \exp\left(-\frac{\mu_i^2}{2\eta_i^2}\right), \\ w_i = \frac{u_i}{u_i + (1-w)}. \end{cases} \quad (37)$$

The distribution in (35) is a Bernoulli-truncated Gaussian distribution with hidden mean  $\mu_i$  and hidden variance  $\eta_i^2$ .

<sup>6</sup>It can be noticed that, for deblurring applications,  $\mathbf{h}_i$  is also the  $i$ th column of the matrix  $\mathbf{H}$  introduced in (2).

APPENDIX B

FAST RECURSIVE COMPUTATIONS  
FOR SIMULATING ACCORDING TO  $f(\mathbf{x} | w, a, \sigma^2, \mathbf{y})$

In the Gibbs sampling strategy presented in Section IV, the main computationally expensive task is the generation of samples distributed according to  $f(x_i | w, a, \sigma^2, \mathbf{x}_{-i}, \mathbf{y})$ . Indeed, the evaluation of the hidden mean and hidden variance in (36) of the Bernoulli-truncated Gaussian distribution may be really costly, especially when the bilinear application  $T(\cdot, \cdot)$  is not easily computable. In this appendix, an appropriate recursive strategy is proposed to make this Gibbs update faster. More precisely, we describe how to update efficiently the coordinate  $i$  of the vector  $\mathbf{x}$  at iteration  $t$  of the Gibbs sampler.

Let  $\mathbf{x}^{(t,i-1)}$  denote the current Monte Carlo state of the unknown vectorized image  $\mathbf{x}$  ( $i = 1, \dots, M$ ):

$$\mathbf{x}^{(t,i-1)} = \left[ x_1^{(t)}, \dots, x_{i-1}^{(t)}, x_i^{(t-1)}, x_{i+1}^{(t-1)}, \dots, x_M^{(t-1)} \right]^T. \quad (38)$$

with, by definition,  $\mathbf{x}^{(t,0)} = \mathbf{x}^{(t-1,M)}$ . Updating  $\mathbf{x}^{(t,i-1)}$  consists of drawing  $x_i^{(t)}$  according to the Bernoulli-truncated Gaussian distribution  $f(x_i | w, a, \sigma^2, \mathbf{x}_{-i}^{(t,i-1)}, \mathbf{y})$  in (21) with:

$$\mathbf{x}_{-i}^{(t,i-1)} = \left[ x_1^{(t)}, \dots, x_{i-1}^{(t)}, x_{i+1}^{(t-1)}, \dots, x_M^{(t-1)} \right]^T. \quad (39)$$

The proposed strategy to simulate efficiently according to (21) is based on the following property.

*Property:* Given the quantity  $T(\boldsymbol{\kappa}, \mathbf{x}^{(0)})$  and the vectors  $\{\mathbf{h}_i\}_{i=1, \dots, M}$ , simulating according to  $f(x_i | w, a, \sigma^2, \mathbf{x}_{-i}^{(t,i)}, \mathbf{y})$  can be performed without resorting to the bilinear application  $T(\cdot, \cdot)$ .

*Proof:* Simulating according to (21) mainly requires to compute the vector  $\mathbf{e}_i$  introduced by (34):

$$\mathbf{e}_i = \mathbf{y} - T(\boldsymbol{\kappa}, \tilde{\mathbf{x}}_i^{(t,i-1)}), \quad (40)$$

with

$$\tilde{\mathbf{x}}_i^{(t,i-1)} = \left[ x_1^{(t)}, \dots, x_{i-1}^{(t)}, 0, x_{i+1}^{(t-1)}, \dots, x_M^{(t-1)} \right]^T. \quad (41)$$

Moreover, by using the decomposition in (31) and by exploiting the linear property of  $T(\boldsymbol{\kappa}, \cdot)$ , the vector  $T(\boldsymbol{\kappa}, \tilde{\mathbf{x}}_i^{(t,i-1)})$  in the right-hand side of (40) can be rewritten as:

$$T(\boldsymbol{\kappa}, \tilde{\mathbf{x}}_i^{(t,i-1)}) = T(\boldsymbol{\kappa}, \mathbf{x}^{(t,i-1)}) - x_i^{(t-1)} \mathbf{h}_i, \quad (42)$$

where  $\mathbf{h}_i$  has been introduced in (34). Consequently, to prove the property, we have to demonstrate that the vector series  $\{T(\boldsymbol{\kappa}, \mathbf{x}^{(t,k)})\}_{k=1, \dots, M}$  can be computed recursively without using  $T(\cdot, \cdot)$ . Assume that  $T(\boldsymbol{\kappa}, \mathbf{x}^{(t,i-1)})$  is available at this stage of the Gibbs sampling and that  $x_i^{(t)}$  has been drawn. The new Monte Carlo state is then:

$$\mathbf{x}^{(t,i)} = \left[ x_1^{(t)}, \dots, x_{i-1}^{(t)}, x_i^{(t)}, x_{i+1}^{(t-1)}, \dots, x_M^{(t-1)} \right]^T. \quad (43)$$

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 Number: 8 Author: vmuser Subject: Replacement Text Date: 9/17/2008 1:22:06 PM function

Similarly to (42), the vector  $T(\boldsymbol{\kappa}, \mathbf{x}^{(t,i)})$  can be decomposed as follows:

$$T(\boldsymbol{\kappa}, \mathbf{x}^{(t,i)}) = T(\boldsymbol{\kappa}, \tilde{\mathbf{x}}_i^{(t,i-1)}) + x_i^{(t)} \mathbf{h}_i. \quad (44)$$

Therefore, combining (42) and (44) allow one to state:

$$T(\boldsymbol{\kappa}, \mathbf{x}^{(t,i)}) = T(\boldsymbol{\kappa}, \mathbf{x}^{(t,i-1)}) + (x_i^{(t)} - x_i^{(t-1)}) \mathbf{h}_i.$$

As a conclusion, the bilinear application  $T(\cdot, \cdot)$  can only be used at the very beginning of the algorithm to evaluate  $T(\boldsymbol{\kappa}, \mathbf{x}^{(0)})$  and the vectors  $\{\mathbf{h}_i\}_{i=1, \dots, M}$ . The resulting simulation scheme corresponding to step 3 of Algorithm 1 is detailed in Algorithm 2.

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#### ALGORITHM 2:

##### Efficient simulation according to $f(\mathbf{x} | w, a, \sigma^2, \mathbf{y})$

For  $i = 1, \dots, M$ , update the  $i$ th coordinate of the vector

$$\mathbf{x}^{(t,i-1)} = [x_1^{(t)}, \dots, x_{i-1}^{(t)}, x_i^{(t-1)}, x_{i+1}^{(t-1)}, \dots, x_M^{(t-1)}]^T$$

via the following steps:

1. compute  $\|\mathbf{h}_i\|^2$ ,
  2. set  $T(\boldsymbol{\kappa}, \tilde{\mathbf{x}}_i^{(t,i-1)}) = T(\boldsymbol{\kappa}, \mathbf{x}^{(t,i-1)}) - x_i^{(t-1)} \mathbf{h}_i$ ,
  3. set  $\mathbf{e}_i = \mathbf{x} - T(\boldsymbol{\kappa}, \tilde{\mathbf{x}}_i^{(t,i-1)})$ ,
  4. compute  $\mu_i$ ,  $\eta_i^2$  and  $w_i$  as defined in (36) and (37),
  5. draw  $x_i^{(t)}$  according to (21),
  6. set  $\mathbf{x}^{(t,i)} = [x_1^{(t)}, \dots, x_{i-1}^{(t)}, x_i^{(t)}, x_{i+1}^{(t-1)}, \dots, x_M^{(t-1)}]^T$ ,
  7. set  $T(\boldsymbol{\kappa}, \mathbf{x}^{(t,i)}) = T(\boldsymbol{\kappa}, \tilde{\mathbf{x}}_i^{(t,i-1)}) + x_i^{(t)} \mathbf{h}_i$ .
- 
- 

#### APPENDIX C

##### SIMULATION ACCORDING TO A BERNOULLI-TRUNCATED GAUSSIAN DISTRIBUTION

This appendix presents a general scheme to generate random variables distributed according to a Bernoulli-truncated Gaussian distribution with parameters  $(w, m, s^2)$  whose pdf is:

$$f(x | \lambda, m, s^2) = (1 - \lambda) \delta(x) + \frac{\lambda}{C(m, s^2)} \exp\left[-\frac{(x - m)^2}{2s^2}\right] \mathbf{1}_{\mathbb{R}_+^*}(x)$$

where  $C(m, s^2)$  has been defined in (23). The generation can be conducted by using an auxiliary binary variable  $\varepsilon$  following the strategy detailed in Algorithm 3.

In the algorithm presented above,  $Ber(\cdot)$  and  $\mathcal{N}^+(\cdot, \cdot)$  denote the Bernoulli and the positive truncated Gaussian distributions respectively. In step 2, the generation of samples distributed according to the truncated Gaussian distribution can be achieved by using an appropriate accept-reject procedure with different instrumental distributions [51]–[53].

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#### ALGORITHM 3:

##### Simulation according to a Bernoulli-truncated Gaussian distribution

1. generate  $\varepsilon$  according to  $\varepsilon \sim Ber(\lambda)$ ,
  2. set  $\begin{cases} x = 0, & \text{if } \varepsilon = 0; \\ x \sim \mathcal{N}^+(m, s^2), & \text{if } \varepsilon = 1. \end{cases}$
- 
- 

#### ACKNOWLEDGEMENTS

The authors would like to thank M. Ting for providing the code to generate point spread functions of MRFM tip and for interesting suggestions regarding this work. The authors are also very grateful to Dr. Dan Rugar who provided the real data used in Section VI as well as a valuable feedback about this paper.

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