Scan Time Optimization for Post-injection PET Scans

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Abstract

Previous methods for optimizing the scan times for PET transmission and emission scans under a total scan time constraint were based on linear non-statistical methods and used noise equivalent counts (NEC) criteria. The scan times determined by NEC analysis may be suboptimal when nonlinear statistical image reconstruction methods are used. For statistical image reconstruction, the predicted variance in selected regions of interest is an appropriate alternative to NEC analysis. We propose a new method for optimizing the relative scan times (fractions) based on analytical approximations to the covariance of images reconstructed by both conventional and penalized-likelihood methods. We perform simulations to compare predicted standard deviations with empirical ones. Results show that for statistical transmission image reconstruction, the optimal fraction of the scan time devoted to transmission scanning is shorter than for conventional transmission smoothing.

I. INTRODUCTION

For PET reconstruction, one has to do two sets of scans, namely transmission and emission scans. One uses the attenuation correction information obtained from the former scan to aid in estimating the radiotracer emission image from the latter one. Conventional methods of reconstruction are based on linear processing of the transmission and emission data, multiplicative correction of attenuation factors in the sinogram domain followed by FBP to reconstruct the emission image. This approach ignores Poisson nature of the data. Recently, there is growing interest on reconstruct/reproject methods for attenuation correction in which one reconstructs the attenuation map and, after possibly some processing in the image domain, this map is reprojected to be used in the attenuation correction factors (ACF) computation. The use of statistical methods for reconstructing attenuation maps as well as emission images is becoming attractive in the medical research community, especially due to faster computers and faster algorithms. In this paper, we reconstruct ACFs using both conventional and penalized-likelihood reconstruct/reproject (PL) methods for post-injection transmission scans. For brevity, we reconstruct emission images with FBP only. Resolution matching is critical in attenuation correction, so we add a post-filtering step to statistical reconstructions to yield approximately Gaussian point spread functions which reduces artifacts from point spread function mismatches. This post-filter reduces the negative sidelobes from the point spread function of penalized-likelihood reconstructions [1].

In this work, we study the effects of emission and transmission scan time duration on the variance of the reconstructed emission image for different reconstruction methods. Particularly we are interested in the optimum scan time fractions under a fixed total scan time constraint, which would result in the smallest variance in a region of interest in the final emission image estimate. Previous studies of scan time optimization [2] were based on NEC criteria with multiple acquisitions of emission and transmission data and focused on conventional reconstructions. The (co)variance approximations developed here might also be useful for other purposes such as determining the weights in a weighted least-squares image reconstruction [3]. We analyze both the conventional and statistical reconstruction cases. We give approximate analytical formulas for conventional and quadratic penalty attenuation map reconstructions. Our analysis is based on Poisson statistics and mathematical approximations [4].

Let $y^{\mathrm{E}} = [y_1^{\mathrm{E}} \dots y_N^{\mathrm{E}}]'$ and $y^{\mathrm{T}} = [y_1^{\mathrm{T}} \dots y_N^{\mathrm{T}}]'$ be emission and post-injection transmission scan count vectors, and let $\mu = [\mu_1 \dots \mu_p]'$ and $\lambda = [\lambda_1 \dots \lambda_p]'$ be attenuation map and emission image pixel value vectors respectively.

We define the survival probabilities as follows: $\alpha_i(\mu) = e^{-l_i(\mu)}$, where $l_i(\mu)$ represents the line integral along projection i of the attenuation map μ . We also define the emission contamination count rate $\kappa_i(\lambda, \mu) = k_i \varepsilon_i \alpha_i(\mu) p_i(\lambda)$. Here k_i is the fraction of emission counts contaminating the transmission data (the portion in the transmission window for rotating rod sources), $p_i(\lambda)$ represents the geometric projections of the true emission image λ , and ε_i contains the detector efficiencies and a scaling factor that accounts for emission scan count rate. We assume that the emission scan measurements $y^{\rm E}$ and the transmission scan measurements with corresponding means:

$$\bar{y}_i^{\mathrm{T}}(\mu,\lambda) = \tau^{\mathrm{T}} \left(b_i \alpha_i(\mu) + r_i^{\mathrm{T}} + \kappa_i(\lambda,\mu) \right), \qquad (1)$$

$$\bar{y}_i^{\rm E}(\lambda,\mu) = \tau^{\rm E} \left(\varepsilon_i \alpha_i(\mu) p_i(\lambda) + r_i^{\rm E} \right).$$
(2)

Here, τ^{T} and τ^{E} are transmission and emission scan times respectively. $l_i(\mu) = [\mathbf{G}\mu]_i \stackrel{\Delta}{=} \sum_{j=1}^p g_{ij}\mu_j$ and $p_i(\lambda) = \sum_{j=1}^p g_{ij}\lambda_j$ are geometric tomographic projections of parameters μ and λ . b_i , r_i^{T} and r_i^{E} are blank scan, transmission scan randoms and emission scan randoms count rates respectively. We assume $\{b_i\}, \{r_i^{\mathrm{T}}\}, \{\varepsilon_i\}, \{r_i^{\mathrm{E}}\}$ and $\{g_{ij}\}$ are known constants throughout this work.

II. ACF ESTIMATION

Attenuation correction is a must for quantitatively accurate emission image reconstruction. We define attenuation correction factors (ACFs) $\gamma_i(\mu) = e^{l_i(\mu)} = 1/\alpha_i(\mu)$. This is the multiplicative factor that corrects for the effects of attenuation in the emission data. We consider two different ways of estimating the ACFs: 1) Conventional smoothing method and 2) Reconstruct/reproject penalized-likelihood (PL) method.

In the non-statistical conventional method, we estimate the emission contamination by:

$$\hat{\kappa}_i = \operatorname{smooth}\left\{k_i \left(y_i^{\mathrm{E}} / \tau^{\mathrm{E}} - r_i^{\mathrm{E}}\right)\right\},\tag{3}$$

and we estimate the ACFs by reciprocating the survival probabilities, that is $\hat{\gamma}_i = 1/\hat{\alpha}_i$, where

$$\hat{\alpha}_i = \operatorname{smooth}\left\{ \left(y_i^{\mathrm{T}} / \tau^{\mathrm{T}} - r_i^{\mathrm{T}} - \hat{\kappa}_i \right) / b_i \right\}.$$
(4)

The smoothing operation is often used to reduce noise in the ACFs. We also use smoothing to reduce noise in the emission contamination estimate in (3).

In a statistical reconstruction, one estimates the ACFs by $\hat{\gamma}_i = e^{l_i(\hat{\mu})}$ where $\hat{\mu}$ is the attenuation map estimate computed by the reconstruction algorithm. The emission contamination estimate (3) is included in the model. The statistical reconstruction is considered in detail in section V.

III. EMISSION IMAGE RECONSTRUCTION

For brevity, we consider here the conventional FBP method to reconstruct emission images. We define the attenuated emission projections function as

$$z_i(\lambda,\mu) = p_i(\lambda)\alpha_i(\mu).$$

A linear unbiased estimate of this function is

$$\hat{z}_i = \operatorname{smooth}\left\{ \left(y_i^{\mathrm{E}} / \tau^{\mathrm{E}} - r_i^{\mathrm{E}} \right) / \varepsilon_i \right\}.$$
(5)

Then, an estimate of the projections $p_i(\lambda)$ can be obtained by:

$$\hat{p}_i = \hat{\gamma}_i \hat{z}_i$$

The emission image is reconstructed by standard FBP method. We use the ramp filter only because the estimate \hat{p}_i is already a smooth estimate of $p_i(\lambda)$. Thus,

$$\lambda = \text{FBP}_{\text{ramp}} \left\{ \hat{p}_i \right\}.$$

IV. EMISSION COVARIANCE ESTIMATES

The covariance of the emission image estimate vector λ obtained by the above procedure can be written as follows:

$$\operatorname{Cov}\left\{\hat{\lambda}\right\} = \boldsymbol{P}\operatorname{Cov}\left\{\hat{p}\right\}\boldsymbol{P}',\tag{6}$$

where the matrix \boldsymbol{P} represents the linear FBP operation with a ramp filter. We need to find the covariance of the random vector $\hat{p} = [\hat{z}_i \hat{\gamma}_i]_{i=1}^N$. The computation of the exact covariance of this expression is computationally intensive and is not desirable. Instead, we prefer to evaluate this covariance as a separable sum of the covariances of the vectors \hat{z} and $\hat{\gamma}$. For this purpose, we consider the Taylor series expansion of $\hat{z}_i \hat{\gamma}_i$ in the neighborhood of $\bar{z}_i \bar{\gamma}_i$ where \bar{z} and $\bar{\gamma}$ are mean values of \hat{z} and $\hat{\gamma}$ respectively. Then:

$$\hat{p}_{i} = \hat{z}_{i}\hat{\gamma}_{i} \approx \bar{z}_{i}\bar{\gamma}_{i} + \bar{\gamma}_{i}(\hat{z}_{i} - \bar{z}_{i}) + \bar{z}_{i}(\hat{\gamma}_{i} - \bar{\gamma}_{i})$$

$$= \bar{z}_{i}\hat{\gamma}_{i} + \bar{\gamma}_{i}(\hat{z}_{i} - \bar{z}_{i}),$$
(8)

and consequently,

$$\operatorname{Cov}\left\{\hat{p}\right\} \approx \mathcal{D}\left\{\bar{z}_{i}\right\} \operatorname{Cov}\left\{\hat{\gamma}\right\} \mathcal{D}\left\{\bar{z}_{i}\right\} + \mathcal{D}\left\{\bar{\gamma}_{i}\right\} \operatorname{Cov}\left\{\hat{z}\right\} \mathcal{D}\left\{\bar{\gamma}_{i}\right\}.$$
(9)

The ACFs $\hat{\gamma}_i$ are not *linearly* related to variables with known covariances. In the conventional method, $\hat{\gamma}_i = 1/\hat{\alpha}_i$. In the statistical method $\hat{\gamma}_i = e^{l_i(\hat{\mu})}$. These are both nonlinear functions. Since the covariance of $\hat{\alpha}$ can be found exactly for the conventional method and the covariance of $\hat{\mu}$ can be approximated for the statistical method, we can linearize these formulas around $\bar{\alpha}_i$ and \bar{l}_i to get an estimate of the covariance of $\hat{\gamma}$. This linearization was the method used in [6] to estimate the variances of the ACFs. But, this linearization is not very accurate for especially the conventional method, because the function f(x) = 1/x cannot be closely approximated by a linear function especially when the denominator (survival probabilities) is close to zero and the variance of the denominator is high.

To overcome this problem, we propose an approximation for the probability distribution function of the ACFs. We assume that $\hat{\gamma}_i$ are lognormal distributed. A random variable is lognormal distributed if its logarithm is normally distributed. We believe this is a very accurate assumption because $\hat{\gamma}_i$ is an estimate of $e^{l_i(\mu)}$ and the projections of any random variable (here $l_i(\hat{\mu})$) can be assumed Gaussian due to the Central Limit Theorem. This provides us extra information about the ACFs. With this assumption, one can compute the mean and variance of $\hat{\gamma}_i$'s directly in terms of mean and variance of $\hat{\alpha}_i$ in the conventional method and in terms of mean and variance of \hat{l}_i in the statistical method.

So, for the conventional method, we get:

$$\bar{\gamma}_i = \frac{\bar{\alpha}_i^2 + \sigma_{\hat{\alpha}_i}^2}{\bar{\alpha}_i^3} \tag{10}$$

and

$$\sigma_{\hat{\gamma}_i}^2 = \frac{(\bar{\alpha}_i^2 + \sigma_{\hat{\alpha}_i}^2)^2 \sigma_{\hat{\alpha}_i}^2}{\bar{\alpha}_i^8}.$$
 (11)

Even with the lognormality assumption, the covariance matrix of $\hat{\gamma}$ is not easy to compute directly. But, the diagonal of the matrix is known. So, we propose this approximation for the covariances:

$$\operatorname{Cov}\left\{\hat{\gamma}_{i},\hat{\gamma}_{j}\right\} \approx \frac{\sigma_{\hat{\gamma}_{i}}\sigma_{\hat{\gamma}_{j}}}{\sigma_{\hat{\alpha}_{i}}\sigma_{\hat{\alpha}_{j}}}\operatorname{Cov}\left\{\hat{\alpha}_{i},\hat{\alpha}_{j}\right\}$$
(12)

$$= \sigma_{\hat{\gamma}_i} \sigma_{\hat{\gamma}_j} \rho(\hat{\alpha}_i, \hat{\alpha}_j), \qquad (13)$$

where $\rho(\hat{\alpha}_i, \hat{\alpha}_j)$ represents the correlation coefficient of the vector $\hat{\alpha}$. In matrix form:

$$\operatorname{Cov}\left\{\hat{\boldsymbol{\gamma}}\right\} \approx \boldsymbol{D}_{1}\operatorname{Cov}\left\{\hat{\boldsymbol{\alpha}}\right\}\boldsymbol{D}_{1}$$

where

$$\boldsymbol{D}_{1} = \mathcal{D}\left\{\frac{\sigma_{\hat{\gamma}_{i}}}{\sigma_{\hat{\alpha}_{i}}}\right\} = \mathcal{D}\left\{\frac{\bar{\alpha}_{i}^{2} + \sigma_{\hat{\alpha}_{i}}^{2}}{\bar{\alpha}_{i}^{4}}\right\}$$

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We make sure that the diagonal of the covariance matrix of $\hat{\gamma}$ matches the variances we get from the lognormal assumption. This formula assumes that the correlation coefficient of $\hat{\gamma}$ is largely determined by the smoothing operator **B** and is the same as the correlation coefficient for $\hat{\alpha}$.

Plugging in the approximation (9) for $\hat{\gamma}$ and writing $\bar{z}_i \approx \bar{\alpha}_i \text{smooth}\{p_i(\lambda^{\text{true}})\}$, we get the following

$$\operatorname{Cov}\left\{\hat{p}\right\} \approx \boldsymbol{D}_{E}\operatorname{Cov}\left\{\hat{z}\right\}\boldsymbol{D}_{E} + \boldsymbol{D}_{T}\operatorname{Cov}\left\{\hat{\alpha}\right\}\boldsymbol{D}_{T}, \qquad (14)$$

where

$$\boldsymbol{D}_{E} \stackrel{\Delta}{=} \mathcal{D} \left\{ \frac{\bar{\alpha}_{i}^{2} + \sigma_{\hat{\alpha}_{i}}^{2}}{\bar{\alpha}_{i}^{3}} \right\}$$

$$\boldsymbol{D}_T \stackrel{\Delta}{=} \boldsymbol{D}_E \mathcal{D} \left\{ \operatorname{smooth} \{ p_i(\lambda^{\operatorname{true}}) \} \right\}.$$

The mean and covariance of \hat{z} can be found easily from the expression (5) since it is linearly related to y^{E} . A simple analysis yields: $\bar{z}_i = \text{smooth}\left\{p_i(\lambda^{\text{true}})e^{-l_i(\mu^{\text{true}})}\right\}$ and from (5) and (2):

$$\operatorname{Cov}\left\{\hat{z}\right\} = \frac{1}{\tau^{\mathrm{E}}} \boldsymbol{B} \mathcal{D}\left\{q_{i}\right\} \boldsymbol{B}', \qquad (15)$$

where $q_i \stackrel{\triangle}{=} (\varepsilon_i \alpha_i (\mu^{\text{true}}) p_i (\lambda^{\text{true}}) + r_i^{\text{E}}) / \varepsilon_i^2$, and **B** is a smoothing convolution matrix along the radial direction of the projection space. It was suggested in [5] that angular smoothing is not desirable in attenuation correction, so we smooth only in radial direction.

For conventional ACF computation, ignoring the noise in the emission contamination estimate, the covariance of $\hat{\alpha}$ can be found from (4) and (1).

$$\operatorname{Cov}\left\{\hat{\alpha}\right\} = \frac{1}{\tau^{\mathrm{T}}} \boldsymbol{B} \mathcal{D}\left\{s_{i}\right\} \boldsymbol{B}^{T}, \qquad (16)$$

where $s_i \stackrel{\Delta}{=} (b_i \alpha_i (\mu^{\text{true}}) + r_i^{\text{T}} + \bar{\kappa}_i)/b_i^2$. Here, **B** is the same smoothing matrix as in (15). The same operator **B** is used to obtain both \hat{z} and $\hat{\alpha}$ to avoid artifacts from resolution mismatch [7, 8]. We used Gaussian smoothing as suggested in [7] which avoids any artifacts in the reconstructed image. The mean of emission contamination can be determined from (3) as $\bar{\kappa} = \mathbf{B} [k_i \varepsilon_i \alpha_i (\mu^{\text{true}}) p_i (\lambda^{\text{true}})]_{i=1}^N$. The variance of $\hat{\alpha}_i$ can be found from (16) as

$$\sigma_{\hat{\alpha}_i}^2 = s_i / \tau^{\mathrm{T}} \sum_k B_{ik}^2.$$

Using (4), one can find the mean values of $\hat{\alpha}$ as

$$\bar{\alpha} = \boldsymbol{B}[\alpha_i(\mu^{\text{true}})]_{i=1}^N.$$
(17)

The variance of the sum over a region of interest in the emission image can be found from (6), (14), (15) and (16) as

$$\operatorname{Cov}\left\{u'\hat{\lambda}\right\} = u'\operatorname{Cov}\left\{\hat{\lambda}\right\}u = \frac{1}{\tau^{\mathrm{E}}}v^{\mathrm{E}} + \frac{1}{\tau^{\mathrm{T}}}v^{\mathrm{T}},\qquad(18)$$

with $v^{\rm E} = \sum_{i=1}^{N} q_i (w_i^{\rm E})^2$ and $v^{\rm T} = \sum_{i=1}^{N} s_i (w_i^{\rm T})^2$, and where u is a vector of ones in the region of interest and zeros elsewhere. We define the vectors

$$w^{\mathrm{E}} \stackrel{\Delta}{=} \boldsymbol{B}' \boldsymbol{D}_{\mathrm{E}} \boldsymbol{P}' u, \ w^{\mathrm{T}} \stackrel{\Delta}{=} \boldsymbol{B}' \boldsymbol{D}_{\mathrm{T}} \boldsymbol{P}' u.$$

V. PENALIZED-LIKELIHOOD ATTENUATION RECONSTRUCTION

While conventional method of ACF computation has been used for some time, reconstruct/reproject methods have gained some interest recently. In a statistical reconstruct/reproject method for ACF computation, an attenuation map estimate $\hat{\mu}$ is found from noisy transmission data by maximizing the penalized-likelihood objective function $\Phi(\mu; y^{T}) = L(\mu; y^{T}) - \beta R(\mu)$, where $L(\mu, y^{T})$ is the log-likelihood function and $R(\mu)$ is a regularizing roughness penalty function. After estimating the attenuation map $\hat{\mu}$, we estimate the ACFs by: $\hat{\gamma}_{i} = e^{l_{i}(\hat{\mu})}$, where $l_{i}(\hat{\mu}) = [G\hat{\mu}]_{i}$ is the geometric projection of the attenuation map estimate $\hat{\mu}$. If one uses FBP for emission reconstruction, then \hat{z} should be smoothed to yield similar resolution with the $\hat{\gamma}$ [9] in order to reduce resolution mismatch artifacts.

A. Resolution

Penalized likelihood (PL) or penalized weighted least squares (PWLS) methods are very attractive image reconstruction methods due to their superb noise reduction properties. The variance weighting in PWLS method reduces the variance of the estimates as compared to penalized unweighted least squares (PULS) or FBP reconstructions, because it makes use of the statistical information in the However, attenuation maps reconstructed measurements. with PL or PWLS methods have non-uniform resolution [1] even with a quadratic penalty. This non-uniform resolution is caused by the variance weighting in PWLS (or PL) method and hence does not exist in a PULS reconstruction. Due to this non-uniform resolution, ACF computation by PL method from a real transmission scan causes resolution mismatch between the emission data and reconstructed ACFs. This mismatch reveals itself as artifacts in the final reconstructed emission image.

Fessler's certainty based penalty [9] yields more uniform resolution in terms of the average FWHM of the point spread function over the image. But, it still has non-uniform resolution in that the psf is not circularly symmetric but the level contours look like ellipses whose orientation are image dependent and space-variant. Stayman and Fessler have recently proposed a new modification to the quadratic penalty [10] which yields more circularly symmetric uniform resolution properties. We used this modification in our reconstructions. This modification makes the resolution properties of the PL method close to PULS method. Quadratic PULS method was shown to be essentially equivalent to FBP method with the following constrained leastsquares (CLS) filter defined in spatial frequency domain by (equation (50) in [9])

$$F_p(u;\beta) = \frac{\operatorname{sinc}(ku) / \operatorname{sinc}(u)}{\operatorname{sinc}(ku)^2 + c\beta u^3}, \ u \in [0, 0.5]$$
(19)

where u denotes spatial frequency, k is the ratio of the detector strip width to the pixel size of the system model, and c is a constant dependent on system geometry. This CLS filter has high negative sidelobes in the space domain. The filters that smooth the ACFs and emission data have to be matched. So, the emission data should be blurred with the same filter (19). But, due to high negative sidelobes of filter in (19), after dividing the appropriately blurred emission data to computed survival probabilities from reconstructions, we get artifacts especially for higher blurring amounts (higher β s) around the boundaries of the image. So, we conclude that the results in [7] does only hold for Gaussian smoothing.

To overcome this problem, we first reconstruct a higher resolution image using a smaller β value than desired and then we filter the projections with the following filter:

$$F_2(u) = \frac{F_p(u;\beta)F_g(u;w)}{|F_p(u;\beta)|^2 + 0.1(1 - \cos(2\pi u))}, u \in [0, 0.5],$$

where $F_g(u; w)$ is the desired Gaussian filter with desired FWHM w. Now, the emission data is also filtered with the Gaussian shaped filter $F_g(u; w)$. This approach reduces artifacts and yields acceptable images. The ACF computation in this case is done as follows:

$$\hat{\mu} = \arg \max_{\mu} \Phi(\mu; y^{\mathrm{T}}),$$

$$\hat{l} = \boldsymbol{B}_{2} \boldsymbol{G} \hat{\mu},$$

$$\hat{\gamma}_{i} = e^{\hat{l}_{i}},$$
(20)

where B_2 is the convolution matrix corresponding to $F_2(u)$ above.

B. Covariance Approximations

The covariance formula in (9) is still valid in PL transmission reconstruction. We use the following first order Taylor series expansion for the ACFs:

$$\hat{\gamma} = e^{\hat{l}_i} \approx e^{\bar{l}_i} + e^{\bar{l}_i} (\hat{l}_i - \bar{l}_i), \qquad (21)$$

where $\bar{l} = B_2 G \check{\mu}$ is the mean projection vector where $\check{\mu} = \underset{\mu \ge 0}{\arg \max} \Phi(\mu; \bar{y}^{T})$ is the image reconstructed with noiseless data. $\check{\mu}$ is a very good approximation for the mean of $\hat{\mu}$ [4]. Note that, we do not use the lognormality assumption here, because we believe that the above approximation is accurate enough and lognormal assumption leads to much more computation. From (21) and (20),

$$\operatorname{Cov}\left\{\hat{\gamma}\right\} = \mathcal{D}\left\{e^{\bar{l}_{i}}\right\} \boldsymbol{B}_{2}\boldsymbol{G}\operatorname{Cov}\left\{\hat{\mu}\right\} \boldsymbol{G}^{\prime}\boldsymbol{B}_{2}^{\prime}\mathcal{D}\left\{e^{\bar{l}_{i}}\right\}.$$

To find the covariance of the implicitly defined estimator $\hat{\mu}$, we use the formulas introduced in [4].

The general form of penalized-likelihood estimates is $\hat{\mu} = \arg \max_{\mu} \Phi(\mu; y^{\mathrm{T}})$, where μ is the parameter vector and y^{T} is the measurement vector. This defines an implicit function $\hat{\mu} = h(y^{\mathrm{T}})$. A first order Taylor expansion of the equation $\nabla \Phi(\mu; y^{\mathrm{T}}) = 0$ around $(\check{\mu}, \bar{y}^{\mathrm{T}})$ yields the following approximation [4]:

$$\operatorname{Cov}\left\{\hat{\mu}\right\} \approx \boldsymbol{Q}\operatorname{Cov}\left\{\boldsymbol{y}^{\mathrm{T}}\right\}\boldsymbol{Q}^{\prime},\tag{22}$$

where $Q = \left[-\nabla^{20}\Phi(\check{\mu},\bar{y}^{T})\right]^{-1}\nabla^{11}\Phi(\check{\mu},\bar{y}^{T})$. We use this formula to evaluate the covariance of the penalized-likelihood

estimate of the attenuation map μ . We again ignore the noise in the emission contamination estimate and use the mean value for it in our approximations. The formula yields:

$$\operatorname{Cov}\left\{\hat{\mu}\right\} \approx \frac{1}{\tau^{\mathrm{T}}} \boldsymbol{H}^{-1} \boldsymbol{G}' \mathcal{D}\left\{\frac{(b_i e^{-\tilde{l}_i})^2 (b_i e^{-l_i^{true}} + \tilde{r}_i^{\mathrm{T}})}{(b_i e^{-\tilde{l}_i} + \tilde{r}_i^{\mathrm{T}})^2}\right\} \boldsymbol{G} \boldsymbol{H}^{-1}$$

where

$$\boldsymbol{H} = \boldsymbol{G}' \mathcal{D} \left\{ b_i e^{-\tilde{l}_i} \left(1 - \frac{(\tilde{r}_i^{\mathrm{T}})(b_i e^{l_i^{true}} + \tilde{r}_i^{\mathrm{T}})}{(b_i e^{-\tilde{l}_i} + \tilde{r}_i^{\mathrm{T}})^2} \right) \right\} \boldsymbol{G} + \beta \boldsymbol{R}$$

Here \boldsymbol{R} is the Hessian of the penalty function and includes the modified penalty weights [10] and $\tilde{r}_i^{\mathrm{T}} = r_i^{\mathrm{T}} + \bar{\kappa}_i$.

In this case, the variance of the sum over a region can be predicted with a formula similar to (18). The emission part of the formula is now $w^{\rm E} = \mathbf{B}' \mathcal{D} \left\{ e^{\bar{l}_i} \right\} \mathbf{P}' u$ and q_i remains same. The transmission part changes a lot due to statistical method as term $w^{\rm T}$ should be changed to:

$$w^{\mathrm{T}} = \boldsymbol{G}\boldsymbol{H}^{-1}\boldsymbol{G}'\boldsymbol{B}_{2}'\mathcal{D}\left\{\bar{z}_{i}e^{\bar{l}_{i}}\right\}\boldsymbol{P}'\boldsymbol{u},$$

and the s_i term should be $s_i = \frac{(b_i e^{-\tilde{l}_i})^2 (b_i e^{-l_i^{\text{true}}} + \tilde{r}_i^{\text{T}})}{(b_i e^{-\tilde{l}_i} + \tilde{r}_i^{\text{T}})^2}$. The most computationally intensive part in this computation

The most computationally intensive part in this computation is the part where $\mathbf{H}^{-1}v^*$ should be computed for $v^* = \mathbf{G}' \mathbf{B}'_2 \mathcal{D} \left\{ \bar{z}_i e^{\bar{l}_i} \right\} \mathbf{P}' u$. This operation can be performed by solving the equation $Hx = v^*$ using iterative methods such as conjugate gradient. Also, we assume the mean for $\hat{\gamma}$ is now, $\bar{\gamma} = e^{\bar{l}_i}$.

These variance predictions are useful, because they do not require hundreds of empirical reconstructions of data [4]. However, they require knowing the true parameters and noiseless sinograms. For real data, these are not known, but one can still get a good approximation of variances by replacing the true parameters by their noisy counterparts [4].

Finally, the optimal time fraction for the emission scan can be found by minimizing the variance in (18) with respect to the emission scan time when total scan time is fixed. For the QPL method, the simple analysis yields

$$\tau_{opt}^{E} = \tau^{total} \frac{v^{\mathrm{E}} - \sqrt{v^{\mathrm{E}} v^{\mathrm{T}}}}{v^{\mathrm{E}} - v^{\mathrm{T}}}.$$

Note that for the conventional method, the above formula is invalid because the $v^{\rm T}$ term is not independent from the scan time duration $\tau^{\rm T}$.

VI. RESULTS

We have done a series of simulations to test the proposed variance predictions and to find the optimal scan times under a total scan time constraint. We used two 2-D images corresponding to attenuation map and emission image to generate noisy transmission and emission data with 150000 and 50000 counts per minute respectively. The true images are shown in Figure 1. The transmission scan had 5% randoms

and an emission contamination of 5%. Emission scan had 10% randoms. Randoms rates were assumed to be constant. The total scan time was 20 minutes. To obtain empirical standard deviations, 300 realizations were generated for each scan time distribution. The emission images were reconstructed with FBP with a smoothing filter that yields about 9 mm FWHM resolution in the image domain. ACFs were computed using the conventional and quadratic penalized-likelihood statistical methods. The resolutions for the ACFs were matched for these two methods. The standard deviations of the sums over the heart region in the reconstructed emission images were found empirically and predicted analytically using the derived formulas. The results are shown in Figure 2 as a plot of standard deviation estimates versus emission scan time fraction. The statistical method not only reduces the overall variance, but also yields a larger optimum emission scan time fraction (about 40%) as compared to the conventional method (about 30%). The standard deviation is reduced by about 15-20% in the statistical method as compared to the conventional method. The predictions seem to match the empirical data for the statistical reconstruction, but the predictions for the conventional method seem to underestimate the standard deviations. We conjecture that, the approximations used in deriving the variance formulas causes the mismatch. We are currently working on improving our approximations. Due to highly nonlinear processing of data however, it is likely that there will be some discrepancy between predicted and empirical standard deviation estimates.

VII. CONCLUSION

We presented new approximate formulas for covariances of reconstructed emission images with conventional and statistical ACF computation for post-injection scans. These formulas can be used to predict the variance of the sum over a region of interest in the final reconstructed emission image instead of expensive empirical reconstructions. These formulas can also be used to determine optimal scan times devoted to emission and transmission scans under a total scan time constraint. Results show that, statistical ACF computation not only reduces the overall standard deviation but also yields higher optimum emission scan time fraction than the conventional method. The covariance approximations for statistical method work well for quadratic penalties, but not for non-quadratic penalties. We plan to extend our work to cover non-quadratic penalties in the future.



Figure 1: The attenuation map and emission image used to generate simulation data.



Figure 2: Standard deviation of the sum over the heart region estimates versus emission scan time fraction for conventional and statistical ACF computations.

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